

# Modelling the Quiescent keV Emission from Magnetars

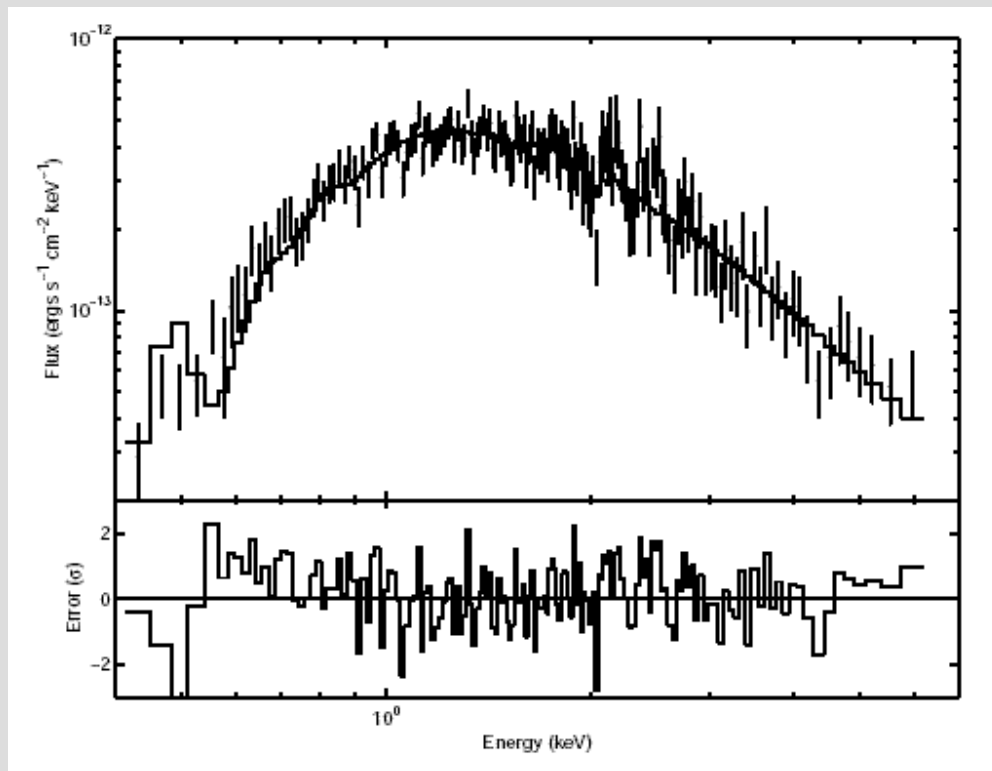
Rodrigo Fernández U. of Toronto

Chris Thompson CITA

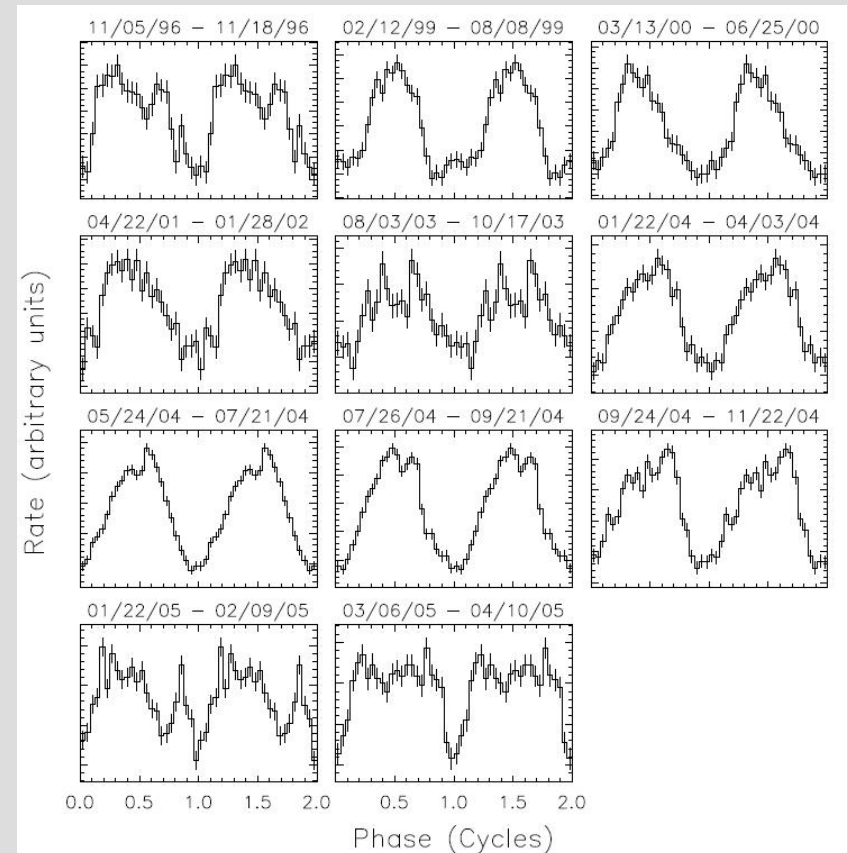
# Outline

- 1) The Model
- 2) Results
- 3) Limitations / Extensions

# Observations (1-10 keV)

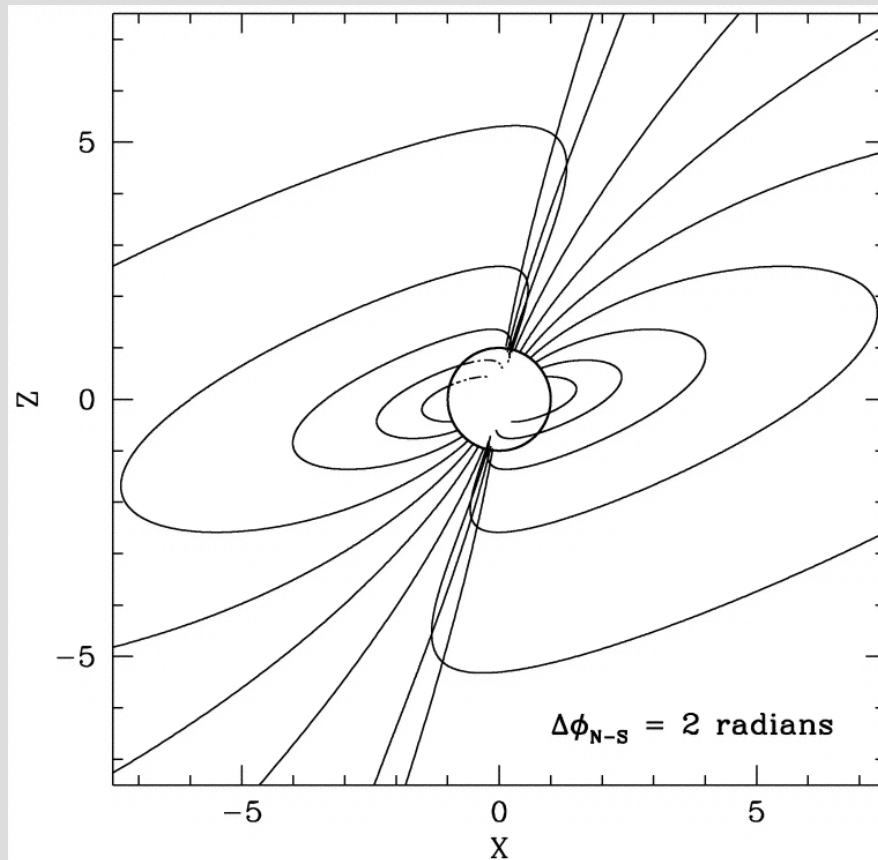


SGR 0526-66 (Kulkarni et al. 2003)



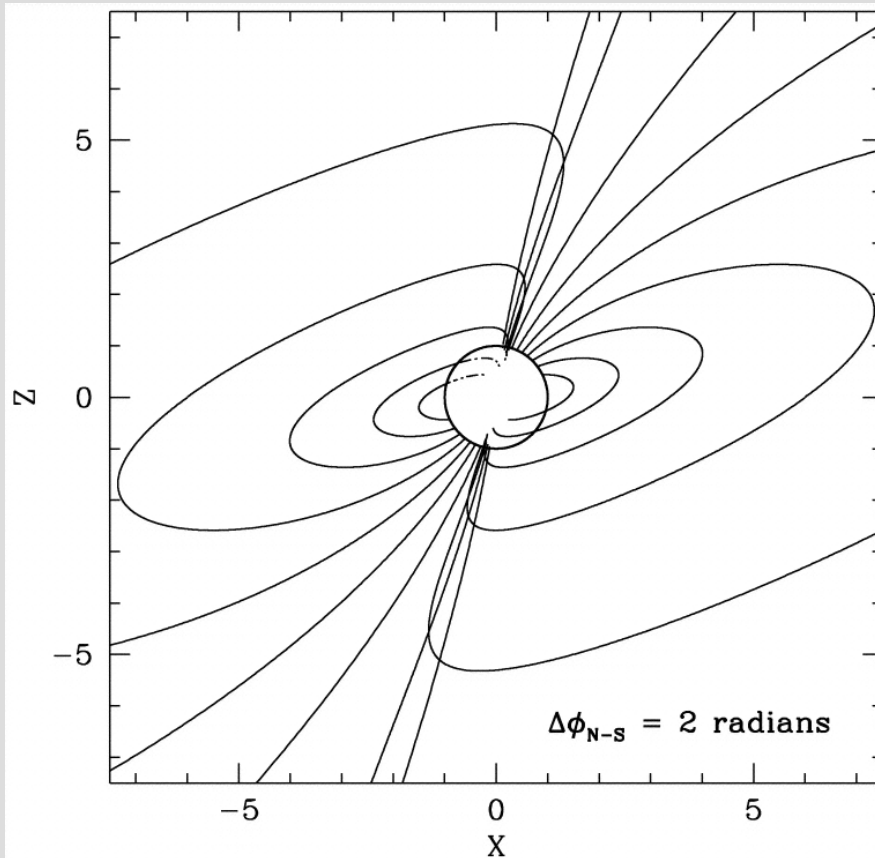
SGR 1806-20 (Woods et al. 2006)

# Model of TLK (2002)



Thompson, Lyutikov, & Kulkarni (2002)

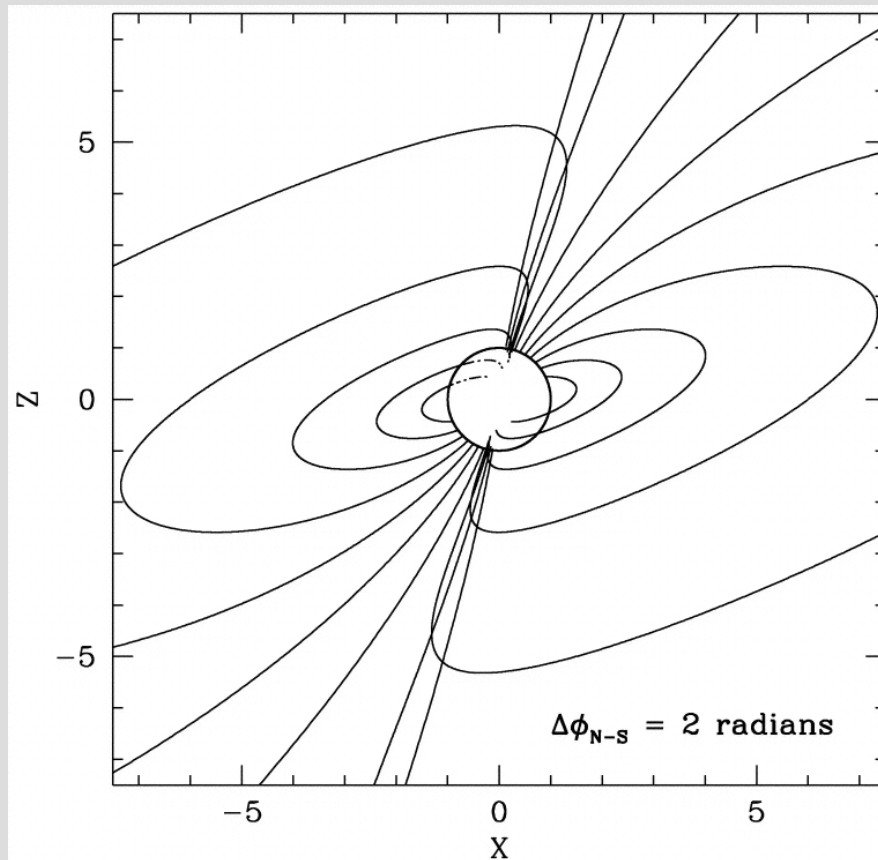
# Model of TLK (2002)



$$\nabla \times \vec{B} = \frac{4\pi}{c} \vec{j}$$

Thompson, Lyutikov, & Kulkarni (2002)

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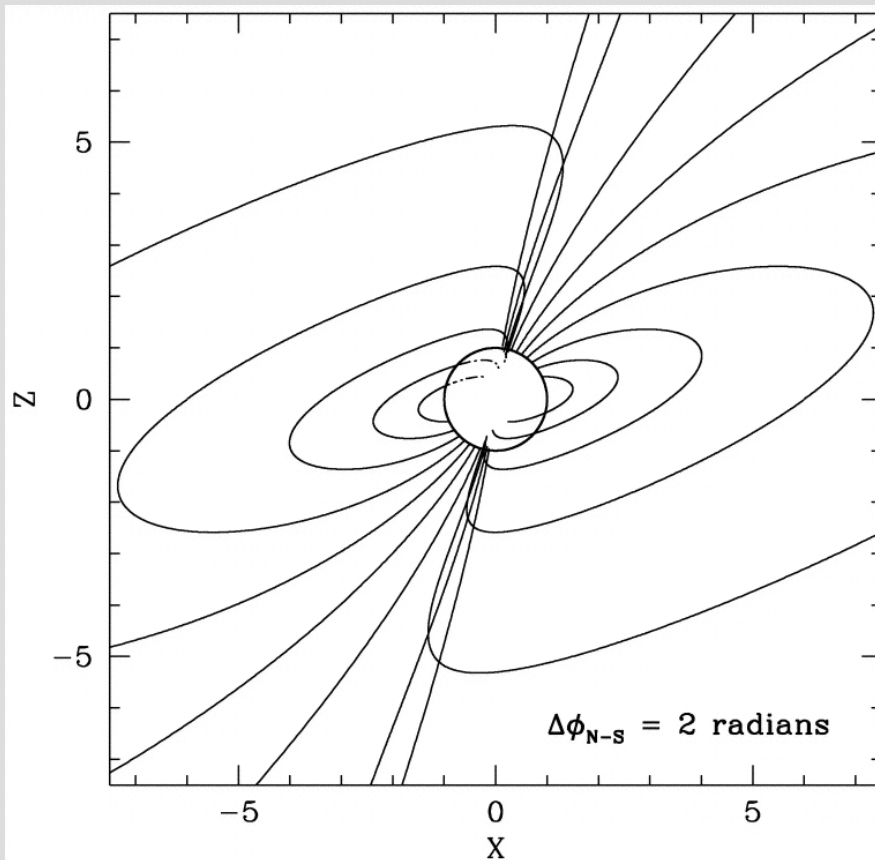


$$\nabla \times \vec{B} = \frac{4\pi}{c} \vec{j}$$

- seed X-rays
  - 1) internal heating
  - 2) surface heating (?)

Thompson, Lyutikov, & Kulkarni (2002)

# Model of TLK (2002)



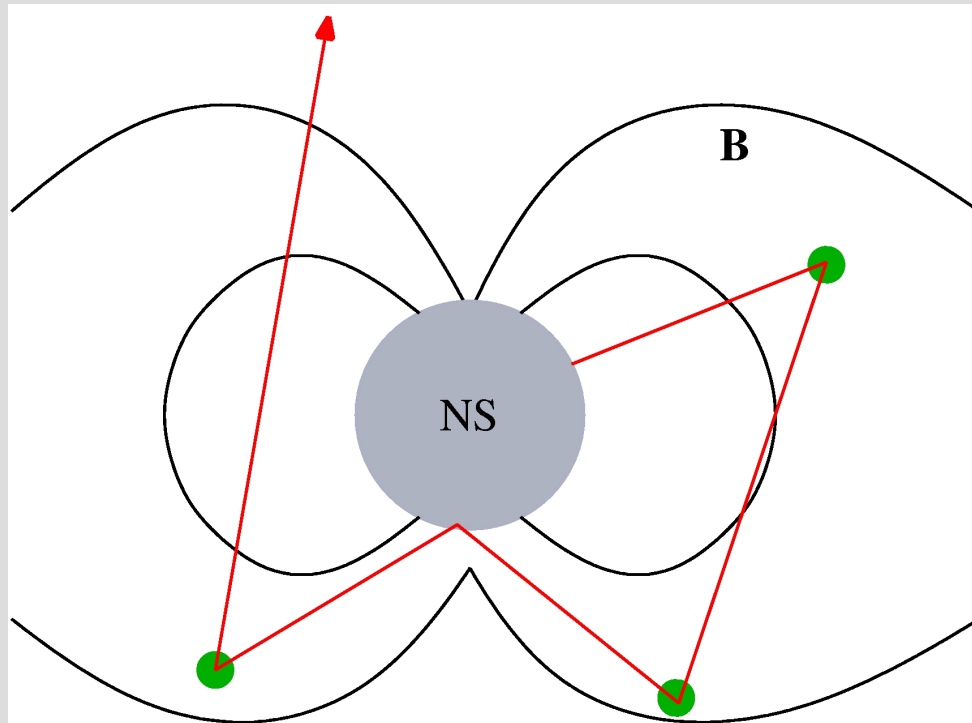
Thompson, Lyutikov, & Kulkarni (2002)

$$\nabla \times \vec{B} = \frac{4\pi}{c} \vec{j}$$

- seed X-rays
  - 1) internal heating
  - 2) surface heating (?)
- resonant cyclotron scattering

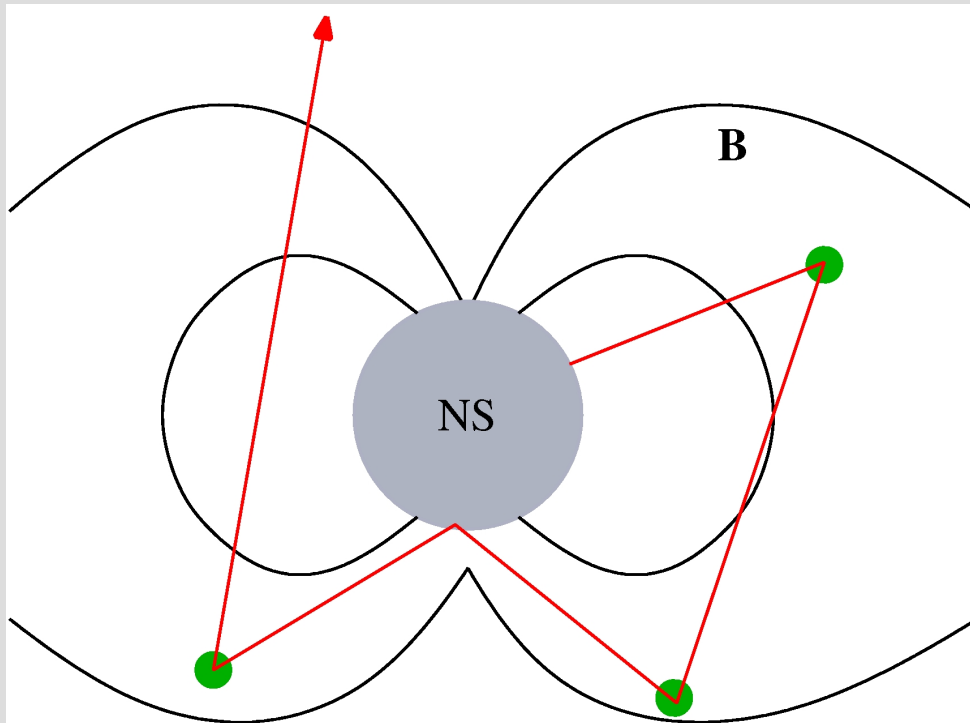
$$\tau_{\text{res}} \sim \frac{\Delta \phi_{\text{N-S}}}{\beta} \sim 1$$

# Monte Carlo Simulation



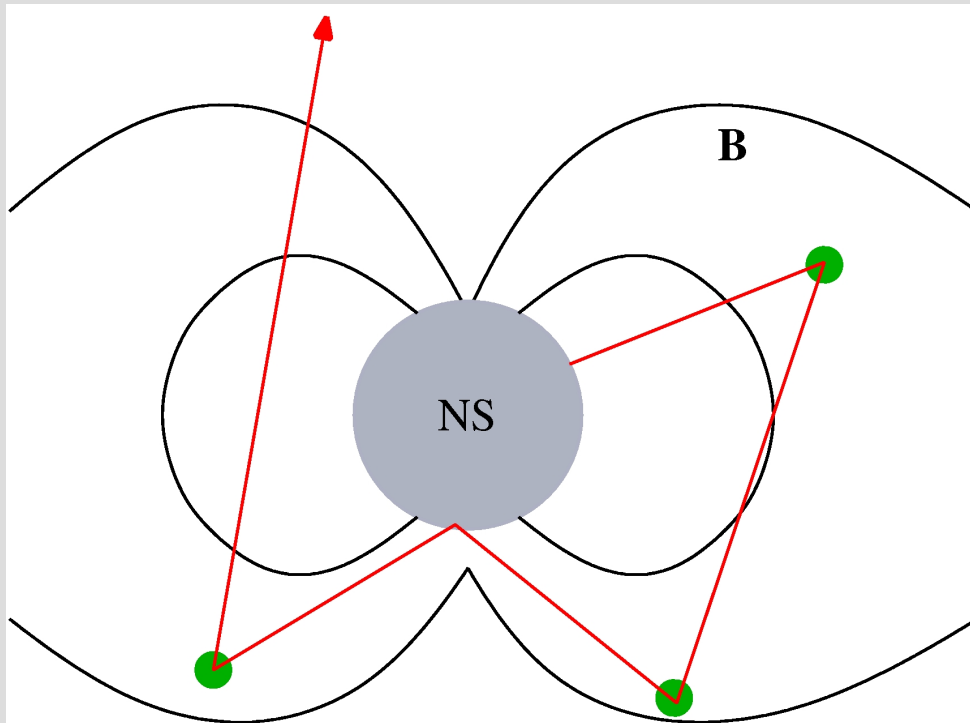


# Monte Carlo Simulation



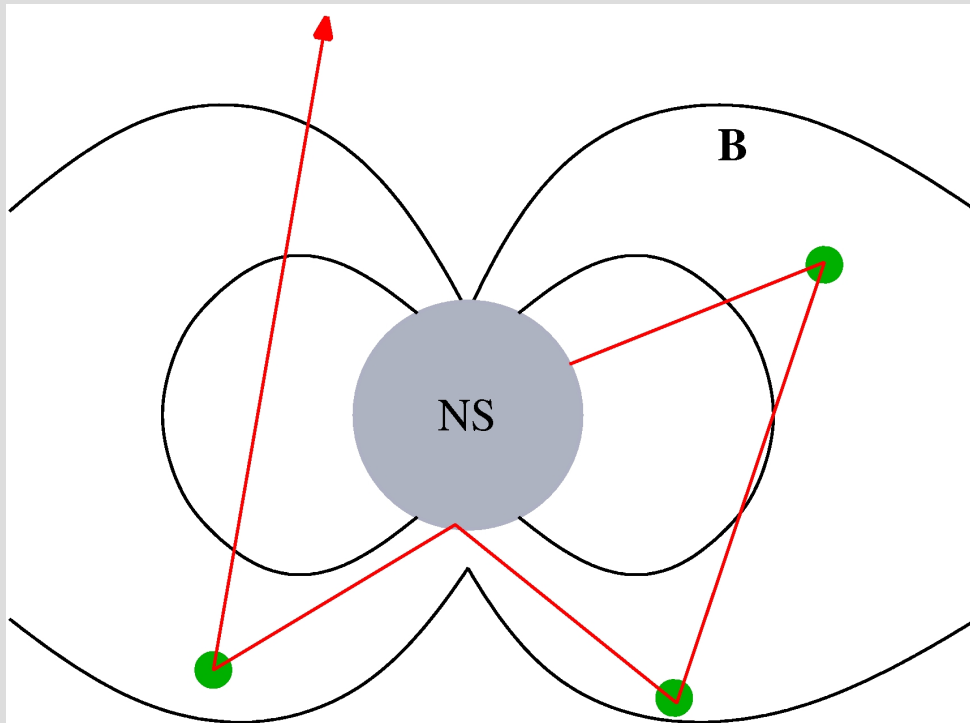
- Multiple scattering

# Monte Carlo Simulation



- Multiple scattering
- 3D, TLK02

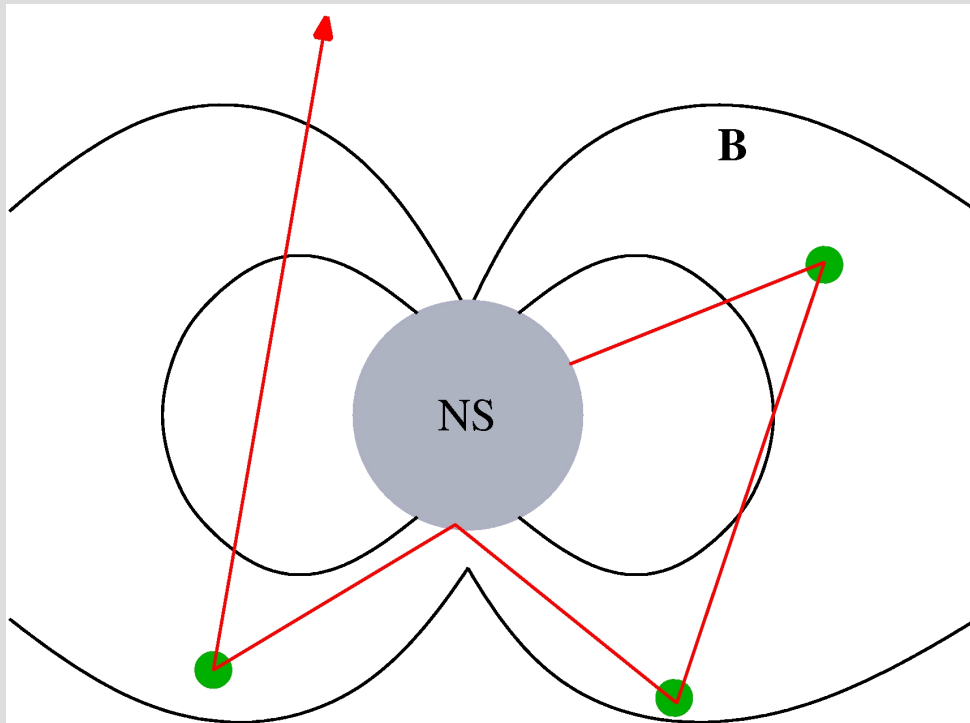
# Monte Carlo Simulation



- Multiple scattering
- 3D, TLK02
- Doppler effect

$$\omega = \frac{\omega_c(r)}{\gamma(1 - \beta\mu)}$$

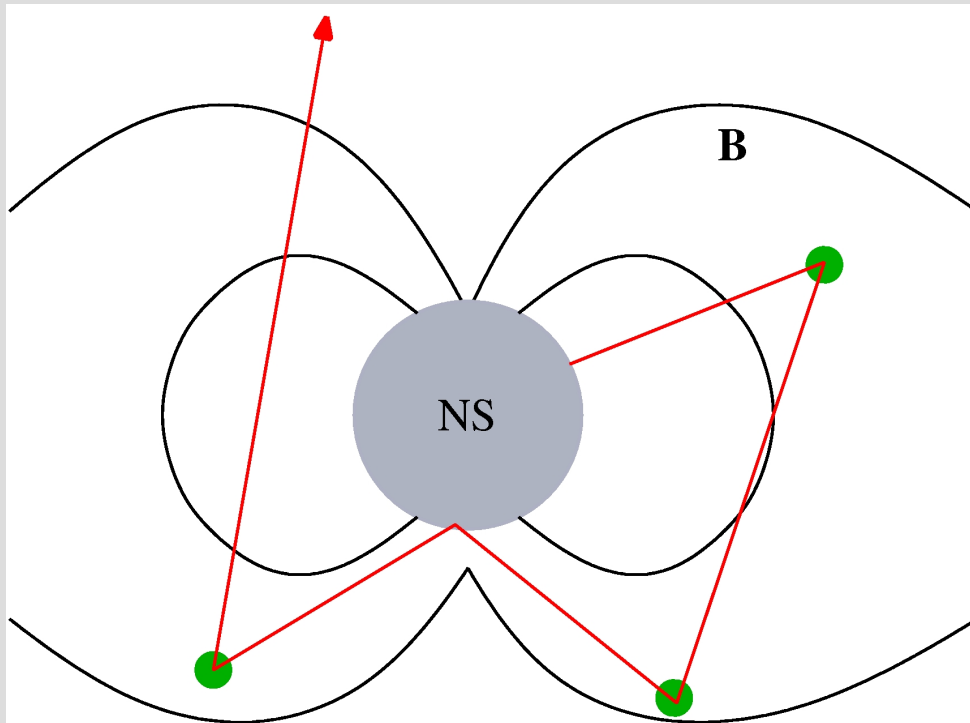
# Monte Carlo Simulation



- Multiple scattering
- 3D, TLK02
- Doppler effect
- polarization

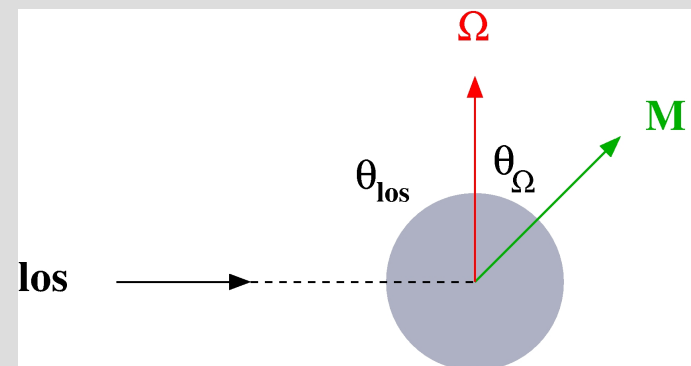
$$\omega = \frac{\omega_c(r)}{\gamma(1 - \beta\mu)}$$

# Monte Carlo Simulation



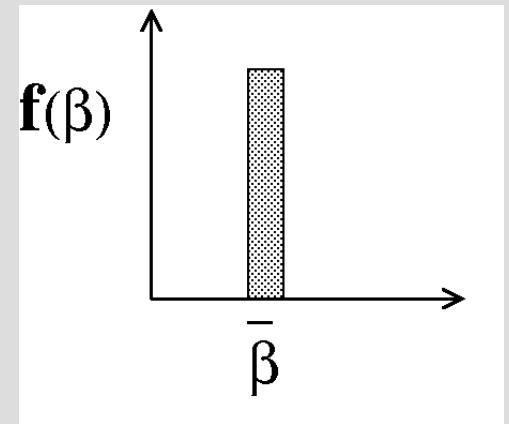
$$\omega = \frac{\omega_c(r)}{\gamma(1 - \beta\mu)}$$

- Multiple scattering
- 3D, TLK02
- Doppler effect
- polarization
- orientation

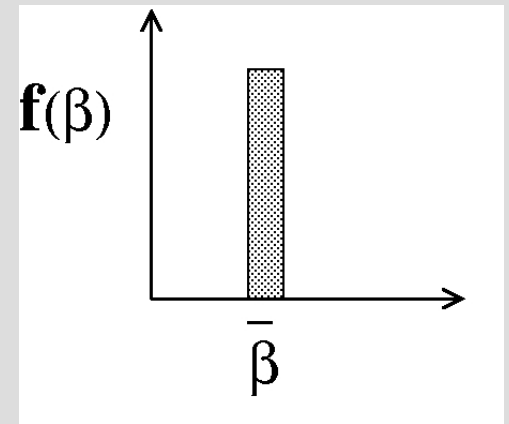
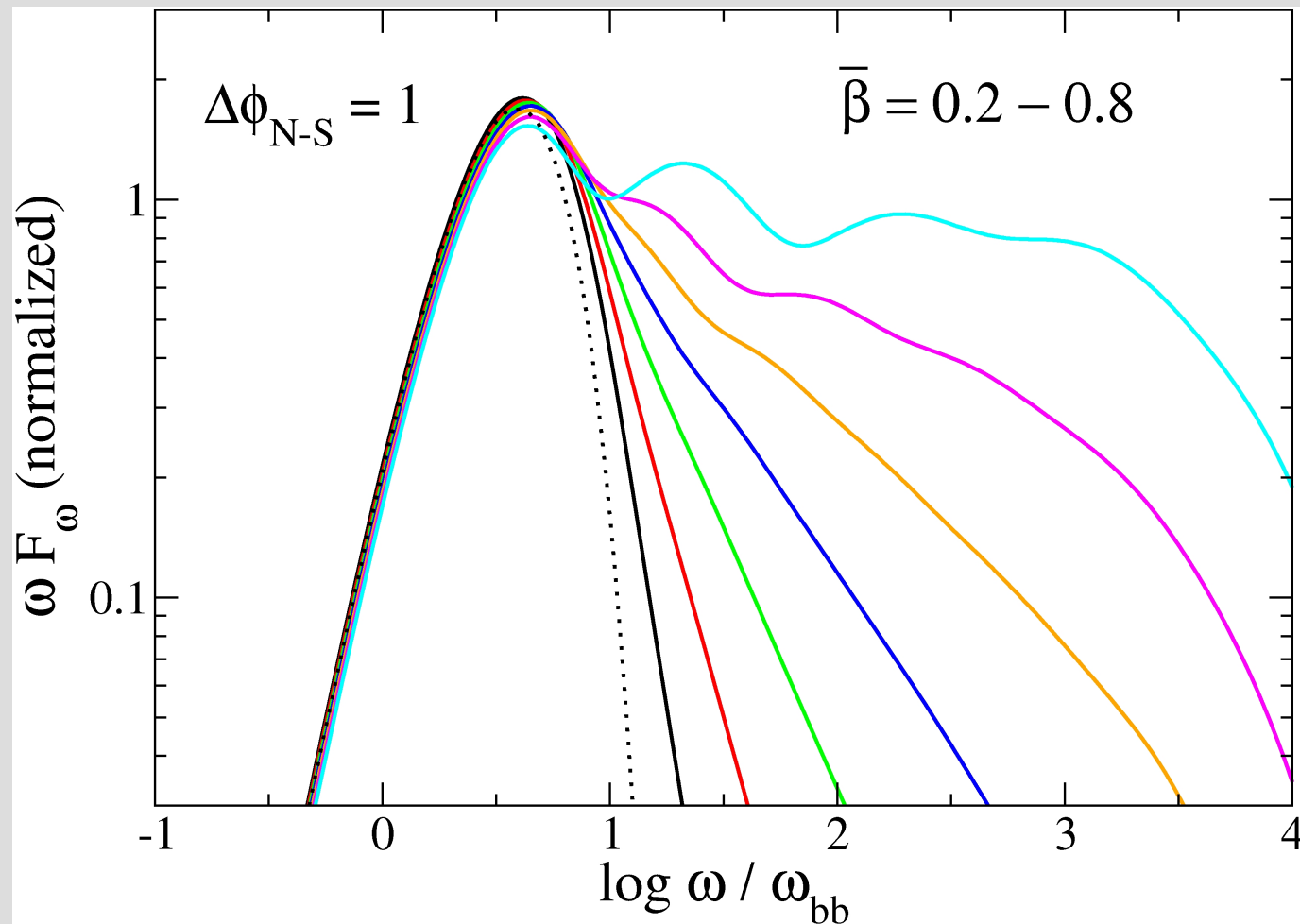


# Results

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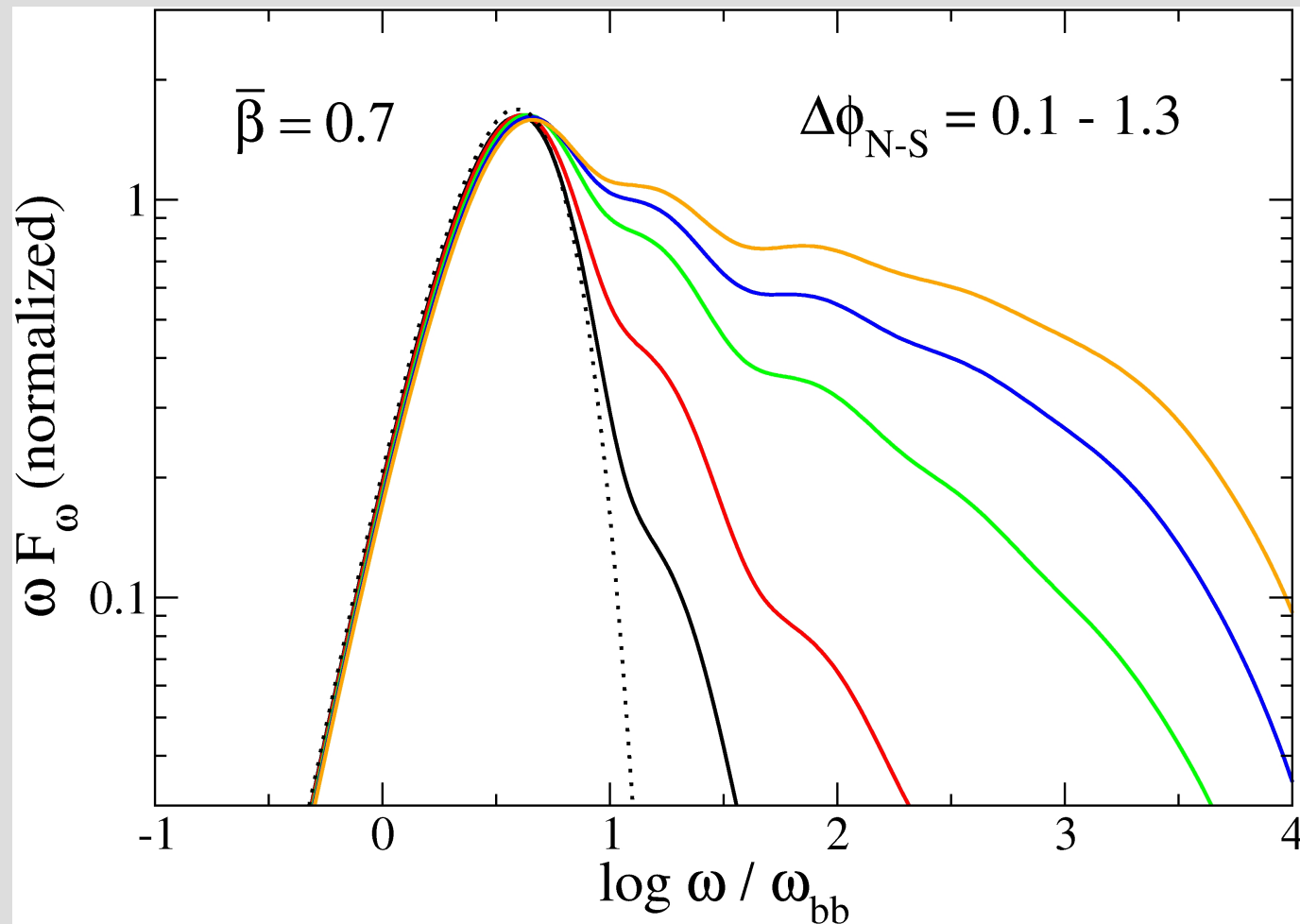
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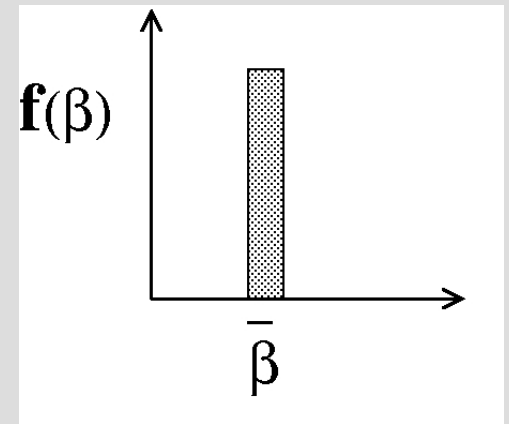
Fernández & Thompson (2006), in prep.



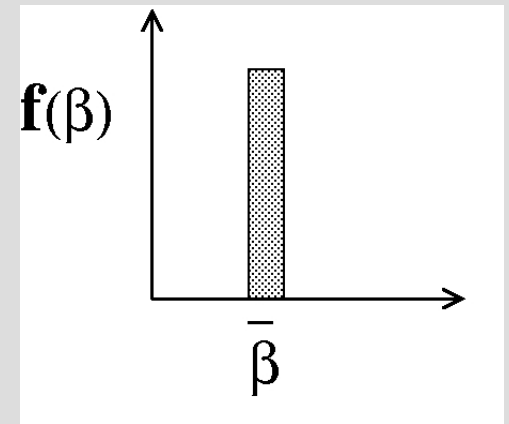
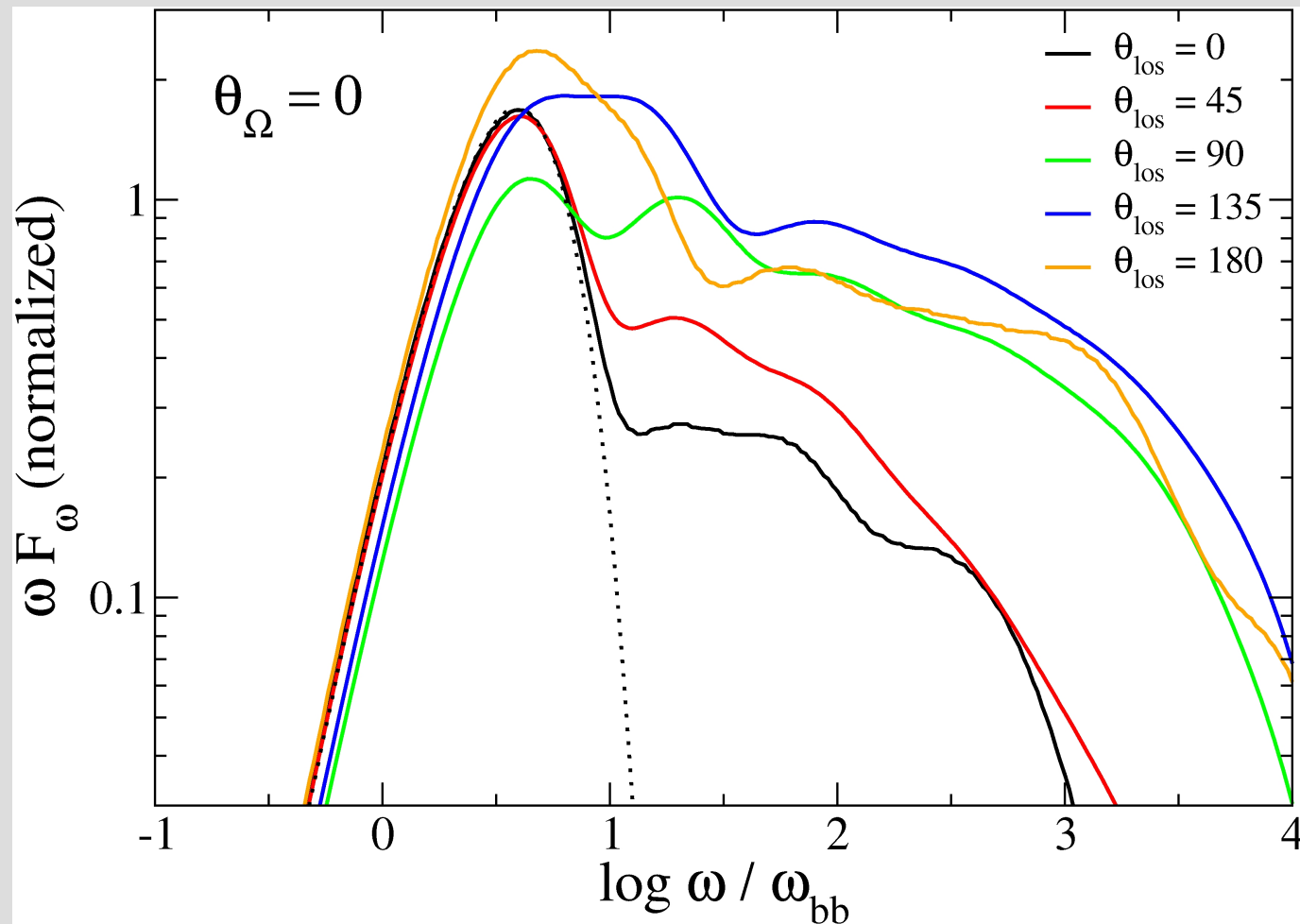
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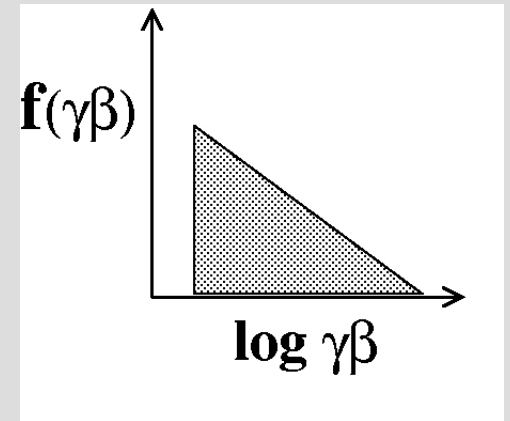


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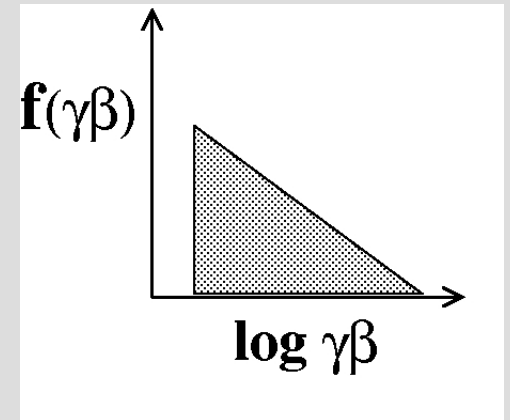
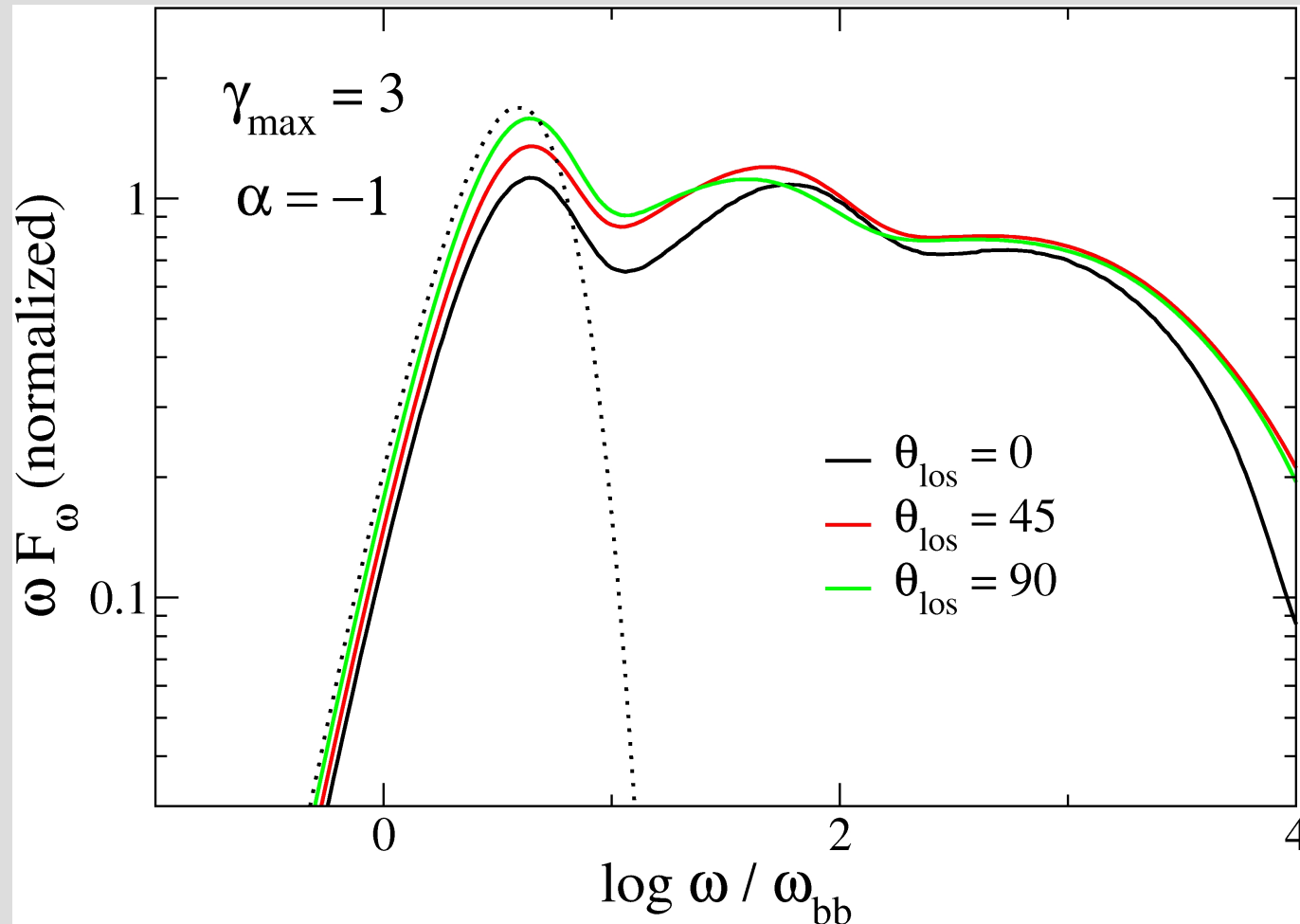
Fernández & Thompson (2006), in prep.

# Results



$$f(\gamma\beta) \propto (\gamma\beta)^\alpha$$

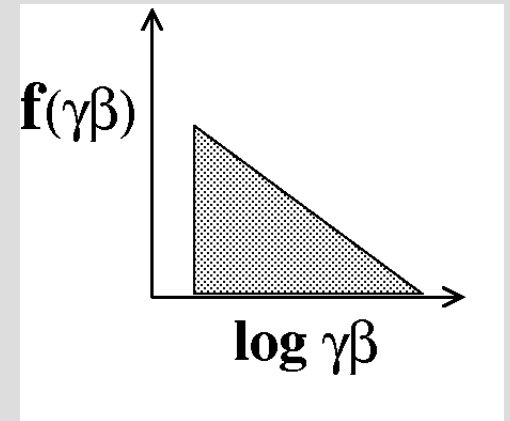
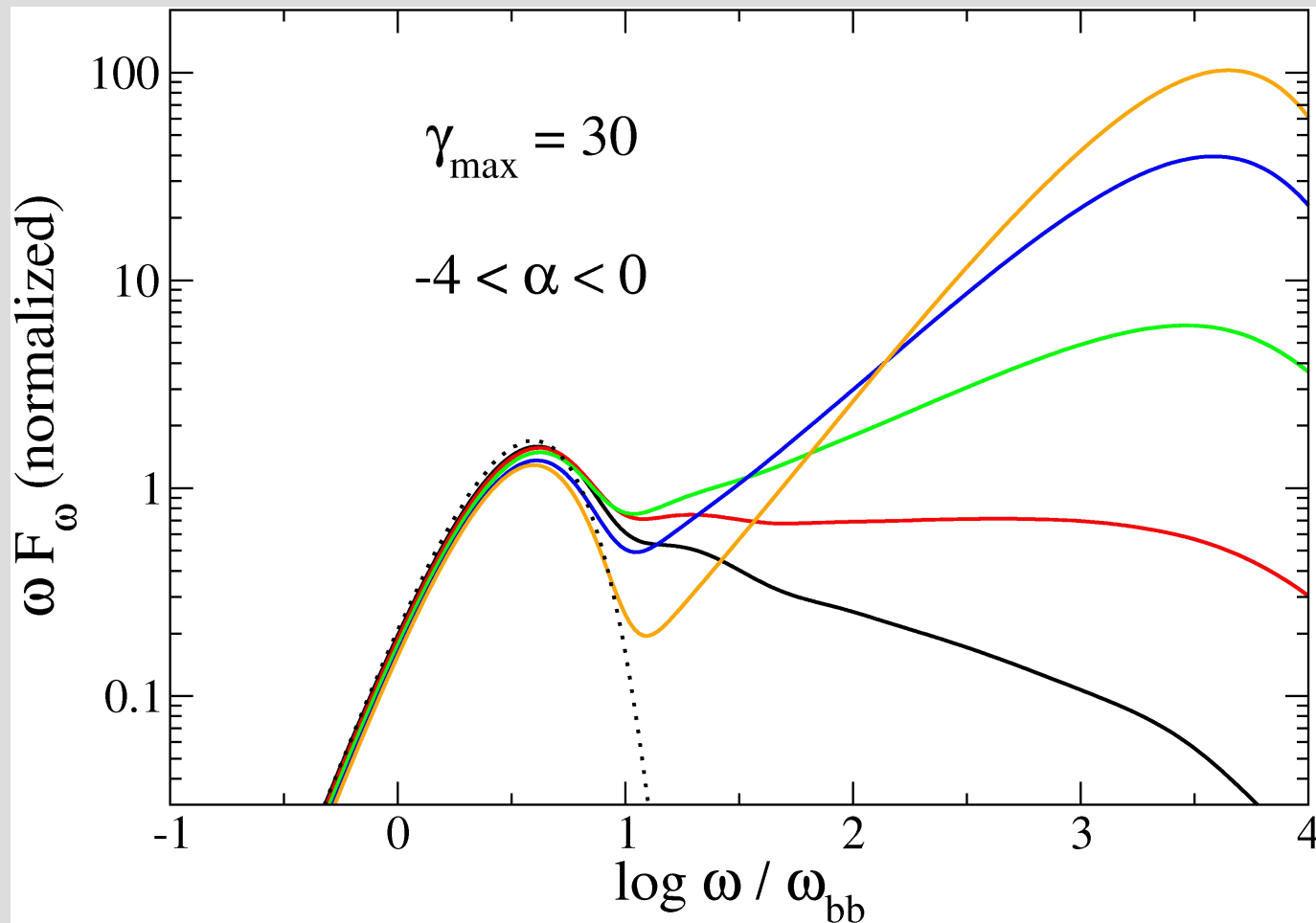
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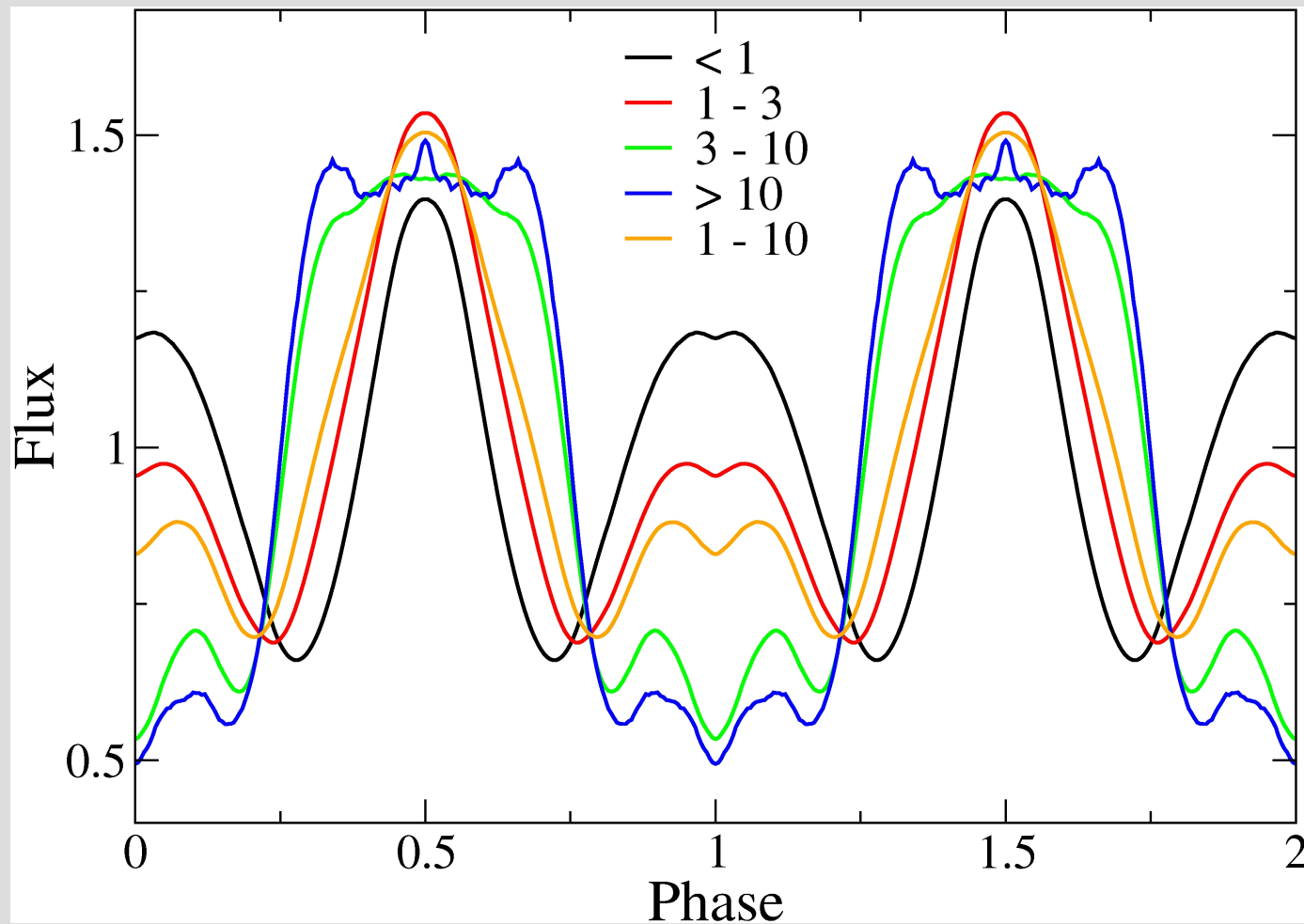
Fernández & Thompson (2006), in prep.

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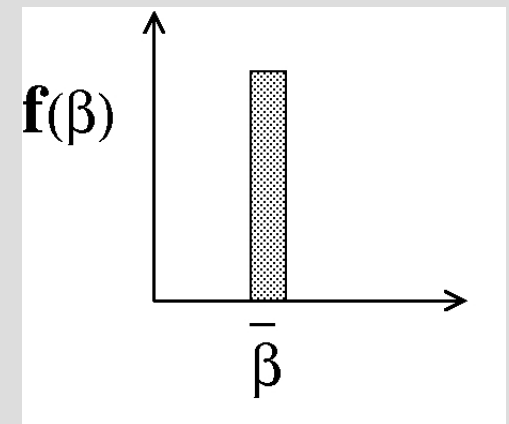


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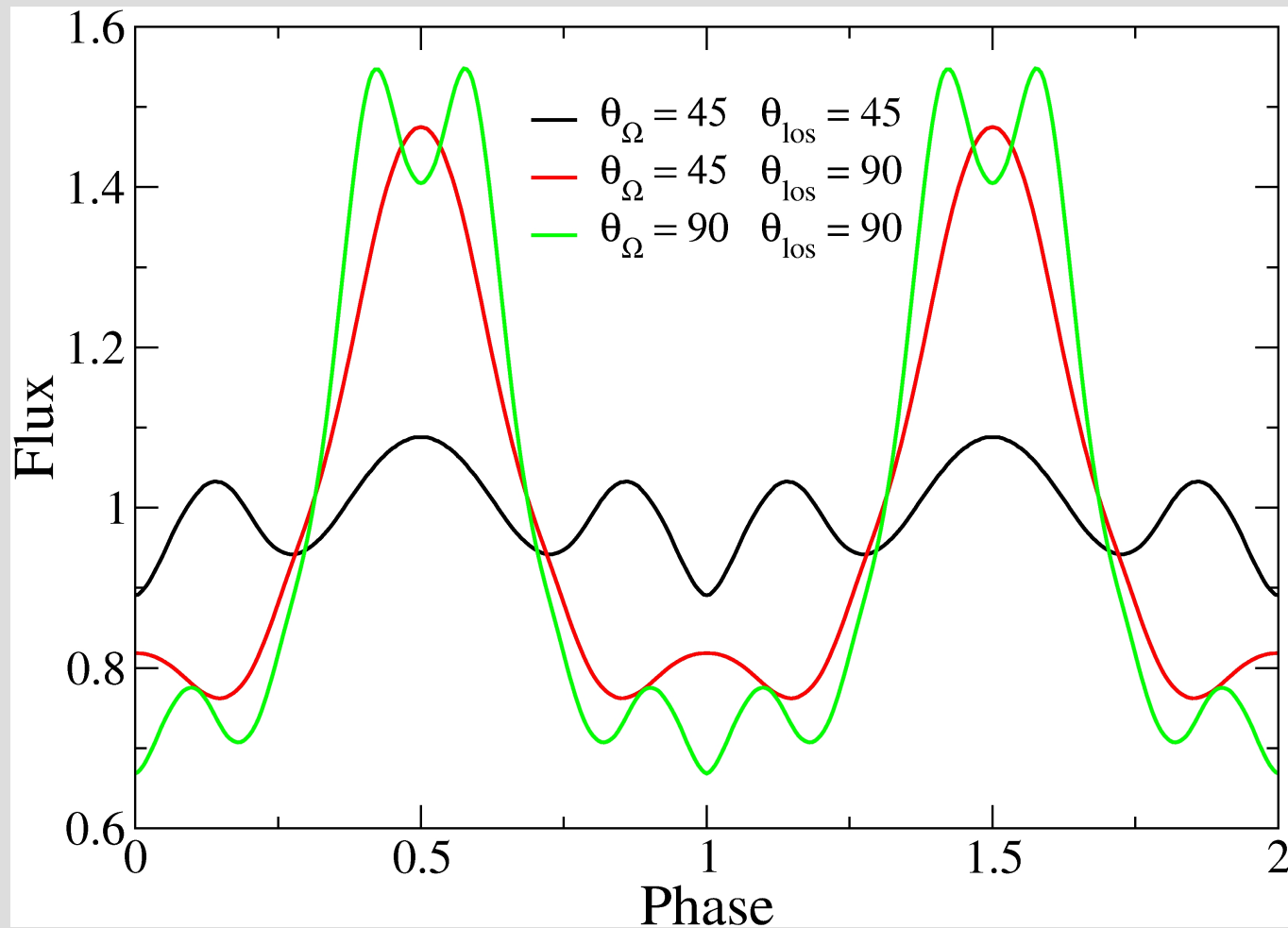
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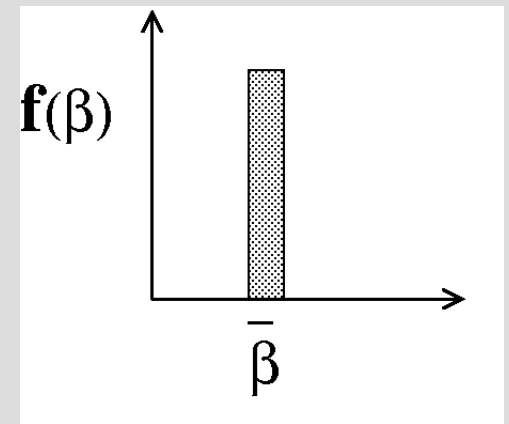
Fernández & Thompson (2006), in prep.



# Results



Fernández & Thompson (2006), in prep.



# Results

- Photon indices:

$$\Gamma \geq 0.5$$



# Results

- Photon indices:

$$\Gamma \geq 0.5$$

- Pulsed fractions from magnetospheric transfer  
( $T_s = \text{constant!}$ )

$$\leq 50 \%$$

# Summary

- Optical depth regulated to  $\sim \Delta\phi_{\text{N-S}}/\beta$  (both modes)
- Spindown-Hardness correlation:  $\Delta\phi_{\text{N-S}}$
- Substantial magnetospheric beaming
- Orientation effects are important

# Limitations

- Sensitive to **uncertain**  $f(\beta, r) \rightarrow$  future work
- Electron recoil
- Light bending
- Inhomogeneous surface T

# Modelling the Quiescent keV emission from Magnetars

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