# Introduction to HI data

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## Where to find general information

#### •Where is STEREO?

- <u>http://stereo.gsfc.nasa.gov/</u> where.shtml
- <u>http://stereo-</u> <u>ssc.nascom.nasa.gov/where/</u>
- •get\_stereo\_sep\_angle
- $\cdot$  stereo\_coord\_info
  - doesn't work with IDLde

📹 solar.bnsc.rl.ac.uk - Solar ssh - SSH !	Secure Shell			>
<u>File E</u> dit <u>V</u> iew <u>W</u> indow <u>H</u> elp				
🖶   🍜 🖻   🎩 🎉   🖻 🖻 🖱	M 🙆 🙆	🎭 🥔 🐶		
👔 Quick Connect 📄 Profiles 👻				
<pre>% Compiled module: GET_SUN. % Compiled module: TIM23D. % Compiled module: ANYTIM2JD. % Compiled module: JULDAY.</pre>				4
% Compiled module: RECPOL. % Compiled module: FILEPATH.				
	STERE0-B	Earth	STERE0-A	
Heliocentric distance (AU)	1.054133	1.010832	0.960081	
Semidiameter (arcsec)	910.348	949.344	999.528	
HCI longitude	155.814	158.418	163.734	
HCI latitude	-3.270	-2.679	-1.988	
Carrington longitude	125.026	127.630	132.946	
Carrington rotation number	2056.653	2056.645	2056.631	
Heliographic (HEEQ) longitude	-2.604	0.000	5.316	
Heliographic (HEEQ) latitude		-2.679	-1.988	
Earth Ecliptic (HEE) longitude	-2.651	0.000	5.356	
Earth Ecliptic (HEE) latitude		-0.000	0.052	
Roll from ecliptic north	-6.122		-0.556	
Roll from solar north	-12.734		-7.516	
Separation angle with Earth Separation angle A with B % Program caused arithmetic err IDL> <mark>-</mark>	or: Floatin	2.666 8.015 g illegal operan	5.357 d	
Connected to solar.bnsc.rl.ac.uk	SSH2 - 3	des-cbc - hmac-md5 - n	one 80x31	



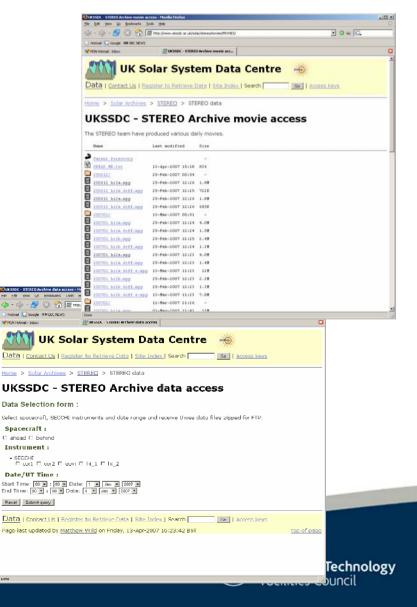
# Where to find HI data/information...

#### $\cdot$ RAL STEREO website

- <u>http://www.stereo.rl.ac.uk/scie</u> <u>nce/</u>
  - $\cdot$  User guide
  - Event List
  - $\cdot$  Gallery

·Daily/Monthly Quicklook Movies

- <u>http://www.ukssdc.ac.uk/solar</u>
   <u>/stereo/movies/MOVIES</u>
- Data
  - <u>http://www.ukssdc.ac.uk/solar</u>
     <u>/stereo/data.html</u>
- ·Email support
  - stereo\_support@ukssdc.ac.uk



# Filename definitions

#### ·20070503\_001000\_s4h1A.fts

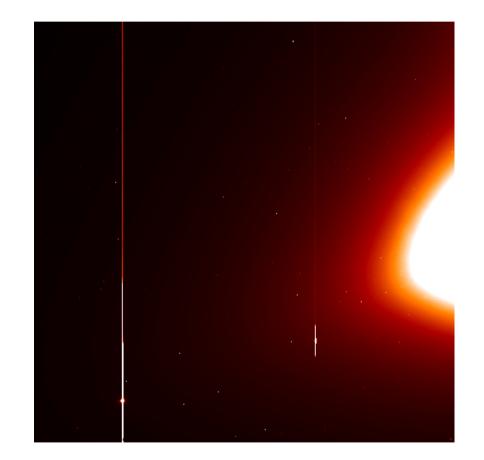
- 20070503 date
- 001000 UT time
- s4 summed image (1024x1024)
- h1 HI 1 image (would be h2 for HI 2 image)
- A STEREO A image (would be B for STEREO B image)
- ·20070503\_000700\_n4h1A.fts
  - n4 hi res (2048x2048) single exposure
  - Only one taken a day will be used for calibration
- ·20070503\_011900\_s7h1A.fts
  - s7 beacon mode image (512x512, 2 hrs)



#### General HI data information

- Images made up of number of exposures
  - HI 1 (nominal values)
    - # of exposures: 30
    - · Exposure time: 40s (1200s total)
    - · Cadence: 40 mins
    - Image size: 1024x1024
  - HI 2 (nominal values)
    - # of exposures: 99
    - Exposure time: 50 seconds (4950s total)
    - · Cadence: 2 hours
    - Image size: 1024x1024
- Exposures summed on board
- Exposures also scrubbed of cosmic rays on board

#### Raw data

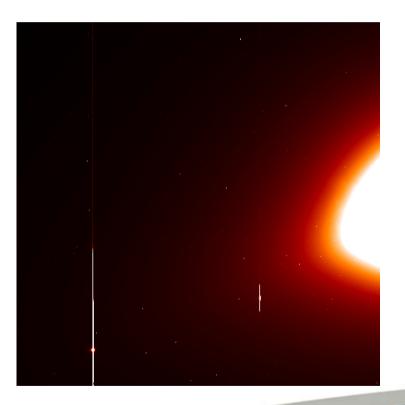




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# secchi\_prep

- $\cdot$  secchi\_prep,files,index,data
- $\cdot$  Files
  - list of all files you want to read
  - Can be mix of EUVI, COR and HI
- $\cdot$ Index
  - Header structures
- Data
  - Data cube
  - If images are all different sizes, then can specify size of data cube that you want returned





# secchi\_prep II

- · HI data
  - Shutterless correction
    - $\cdot$  No shutter
    - Image smeared as readout and CCD cleared out
  - Flatfield
    - Provisional pre-flight flatfield
    - · Calibration on going



# secchi\_prep III

#### Warning

- Early data does not work with secchi\_prep
- All data from beginning of mission will be re-processed at a date TBD
- Number of files that can be read into secchi\_prep depends on machine memory



# Header keywords

 $\cdot$  naxis1/2 – length of 1<sup>st</sup> (columns, x) and 2<sup>nd</sup> (rows, y) axes

· date\_obs - date/time of start of CCD readout

 $\cdot$  exptime – exposure time of all exposure that make up image  $\cdot$  summed – combines summing from CCD and IP to get one number for number of rows and columns being summed on CCD, SEB and ground

- ipsum

- sumrow, sumcol, ccdsum
- detector name of telescope within SECCHI
- obsrvtry name of satellite



# Header keywords II

 $\cdot$  ctype1(a)/2(a) – string representing coordinate axis

- <u>http://stereo-</u>
   <u>ssc.nascom.nasa.gov/coordinates\_explanation.shtml</u>
- Thompson, 2006, A&A, 449, 791
- $\cdot$  cunit1(a)/2(a) units of coordinates

 $\cdot$  crval1(a)/2(a) – reference data coordinates corresponding to centre of image

 $\cdot$  pci\_j(a) – coordinate transformation matrix: rotation information is included

 $\cdot$  cdelt1(a)/2(a) – height/width of pixels

 $\cdot\,pv2\_1$  – parameter which encodes information about optical properties of telescope



# Header keywords III

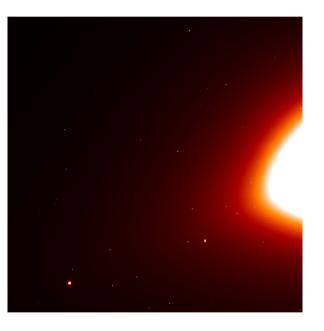
- $\cdot$  More info on header keywords
  - <u>ftp://louis14.nrl.navy.mil/pub/secchi/ssw/doc/FITS\_ke</u> <u>ywords.pdf</u>
- •Warning!
  - Values of header keywords have changed during mission
  - All values will be correct when data is reprocessed
  - Don't hard code header values in to programs until data is reprocessed
  - Need to check that header values are reasonable!



#### Saturated Pixels

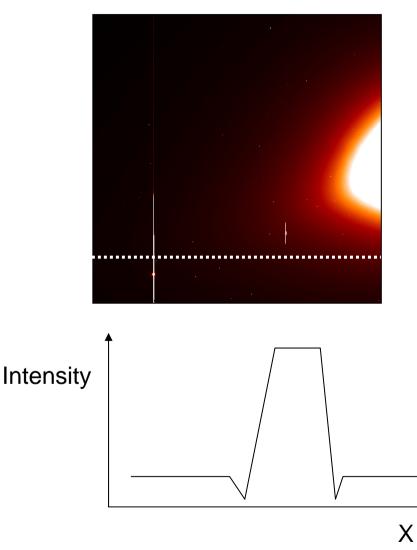
 Brightest objects (e.g. planets & stars) will saturate

- Blooming
  - Occurs up & down column
  - Not side to side like LASCO
- $\cdot$  Can remove blooming





# Saturated Pixels II



• Must deal with saturated columns before doing background subtraction

• Anomolous low values in background

•Anomolous high values in background subtracted data

- Bright stripes in data
- <u>http://www.ukssdc.ac.uk/</u> <u>solar/stereo/movies/MOVI</u> <u>ES/200705\_hi1a.mpg</u>



# Background subtraction

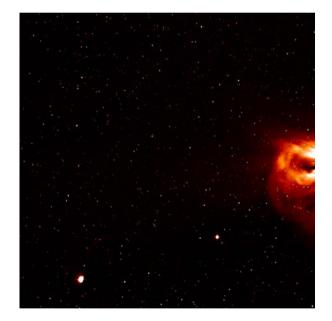
- $\cdot$  bg = min(data,dimension=3)
  - Need to also account for missing blocks and saturated columns when using minimum
- $\cdot$  bg = median(data,dimension=3)
  - More robust
  - But not physically meaningful?

 $\cdot \mbox{Time scale: several days depending on conditions}$ 

•Warning! This will only work if spacecraft is not rolling too much (few degrees) during period of observations

 <u>http://www.ukssdc.ac.uk/solar/stereo</u> /movies/MOVIES/200702\_hi2b.mpg

•Looking into generating monthly backgrounds which could be used for subtraction





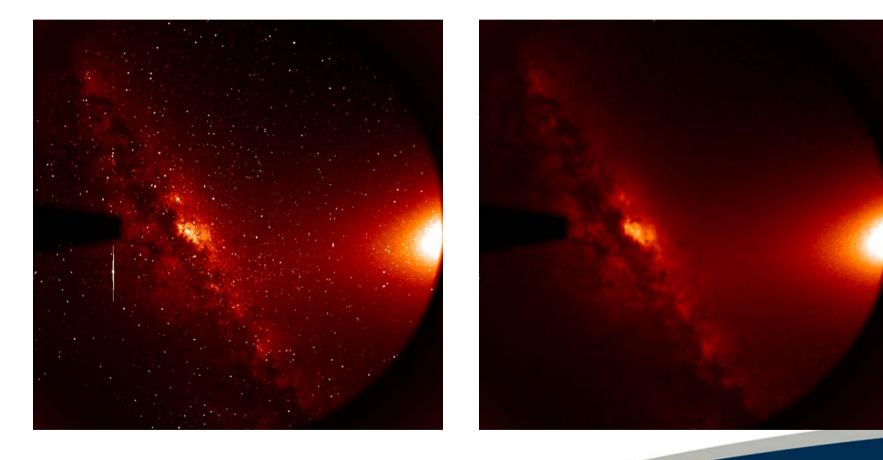
#### Star Removal

Finds peaks in data over a user specified threshold
Uses successive over relaxation method to fill in stars with values from surrounding pixels



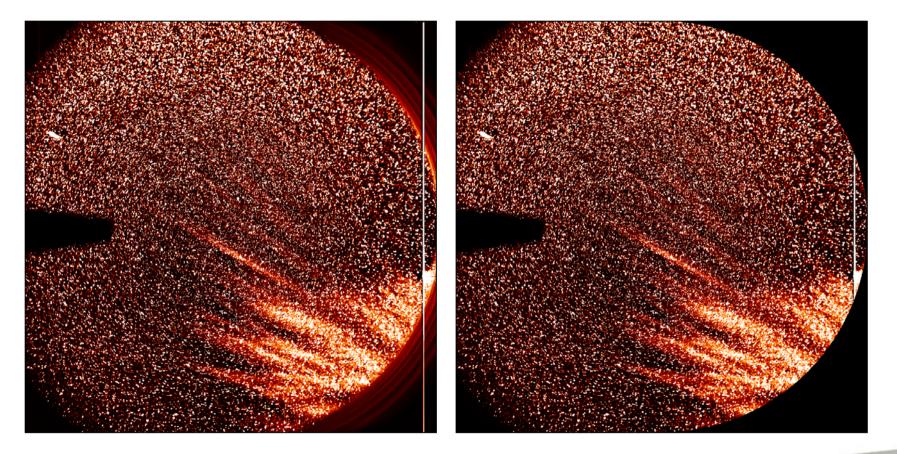


# Star Removal II



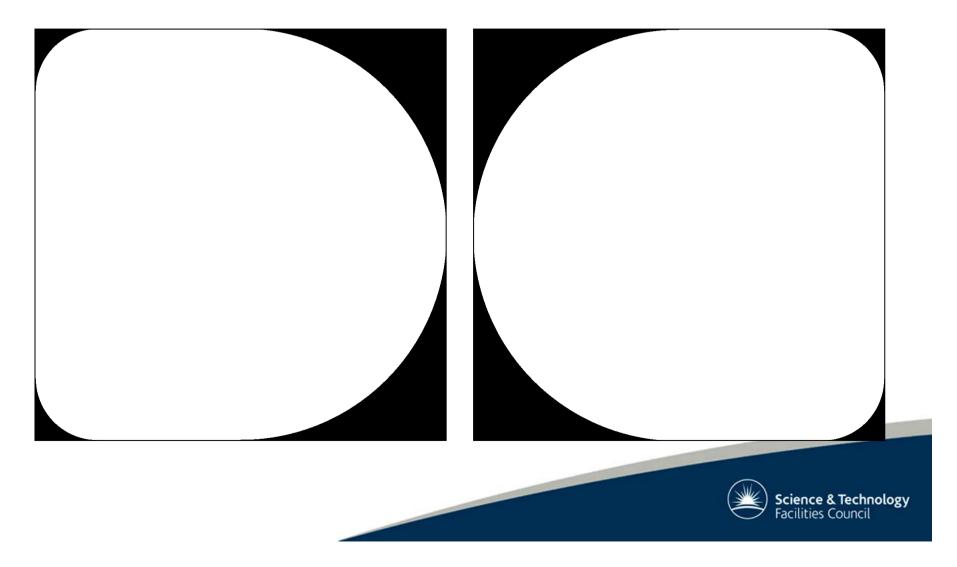


#### HI 2 Mask



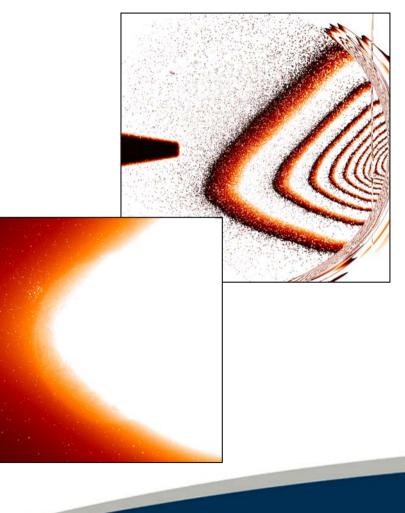


# HI 2 Mask II



# Things to look out for!

- ·High/low word images
  - <u>http://www.ukssdc.ac.uk/solar/</u> <u>stereo/movies/MOVIES/200702</u> <u>\_hi2a.mpg</u>
  - Mucks up background subtraction!
- $\cdot$  Mixed images
  - <u>http://www.ukssdc.ac.uk/solar/</u> <u>stereo/movies/MOVIES/200705</u> <u>\_hila.mpg</u>
  - Mucks up background subtraction!
- $\cdot$ Earth and Moon
  - HI B observations only
  - <u>http://www.ukssdc.ac.uk/solar/</u> <u>stereo/movies/MOVIES/200701</u> <u>\_hi2b.mpg</u>
- Ghosts
  - <u>http://www.ukssdc.ac.uk/solar/</u> <u>stereo/movies/MOVIES/200701</u> <u>\_hi1a.mpg</u>





## Image manipulation

Festival

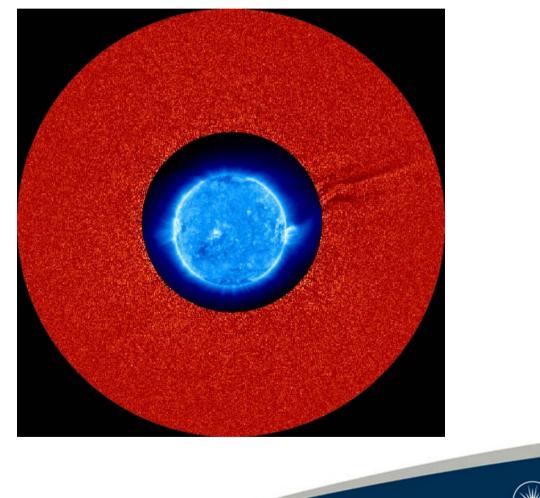
- <u>http://www.ias.u-psud.fr/stereo/festival/</u>

· Solarsoft package

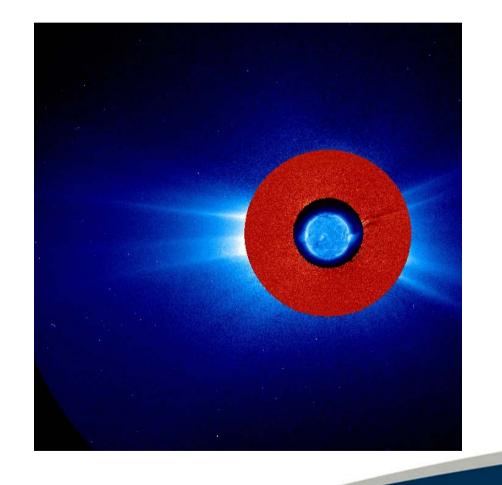
 $\cdot$  'simultaneous, fast and easy manipulation of SECCHI/STEREO, EIT/SOHO and LASCO/SOHO images'



#### Festival Movie – EUVI & COR1



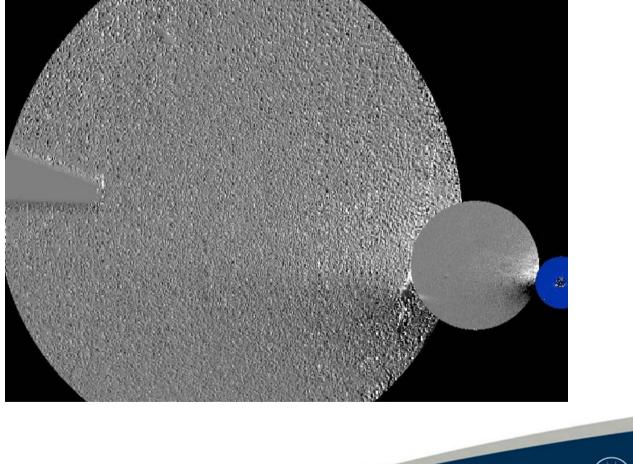
# Festival Movie – EUVI, COR1 & COR2



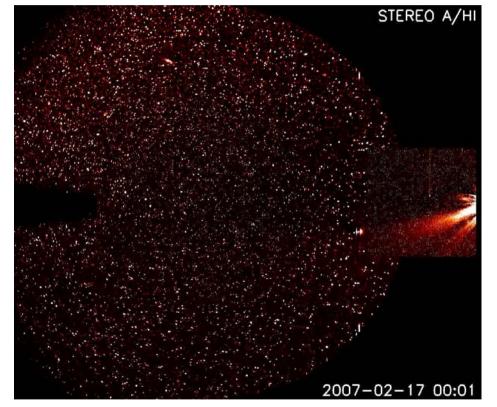


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#### Festival Movie – All SECCHI

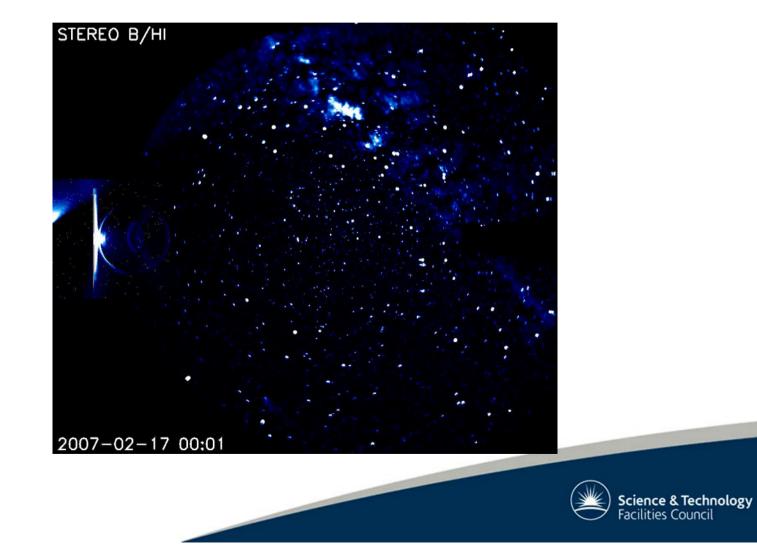


# STEREO A 17/02/07

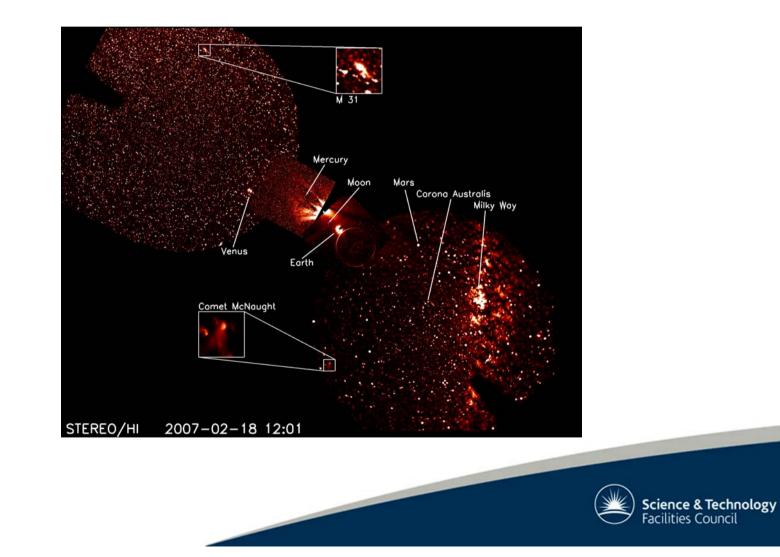




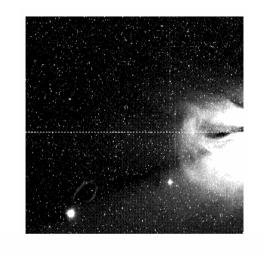
# STEREO B 17/02/07

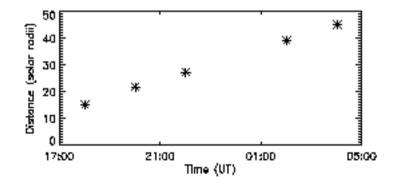


#### STEREO A & B 18/02/2007



# **CME** Velocity





- Example 24/01/07
   RAH manual front detection
- Amount of structure makes it hard to know what to track
- Automatic detection and tracking will be developed

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# **CME** Density

- $\cdot$  Calibration ongoing
  - Absolute intensity
- ·Vourlidas et al. 2000, ApJ, 534, 456
  - Mass calculations from LASCO data
  - Details Thomson scattering assumptions

