

# Good progress in the extrapolation of photospheric magnetograms

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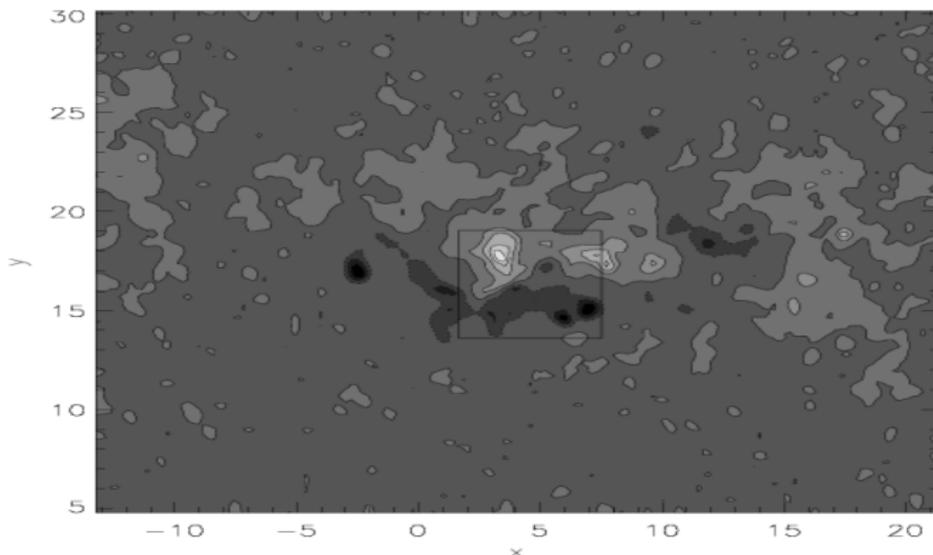
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## A solar example

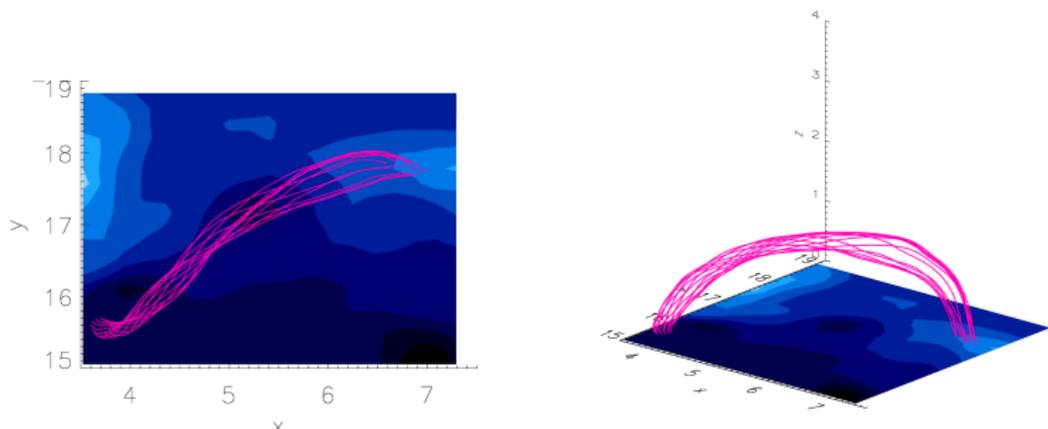
Bastille event: IVM 2000-07-14 16:33  $\subset$  MDI



Extrapolated field:  $\sigma_J = 0.14$        $\max |\vec{\nabla} \cdot \vec{B}| = 1.5 \times 10^{-12}$

$\sigma_J$  is the current-averaged sine angle between  $\vec{B}$  and  $\vec{J}$ .

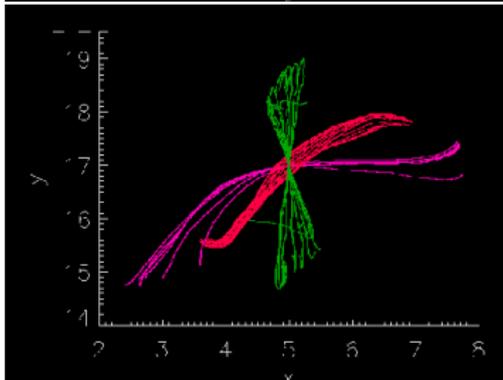
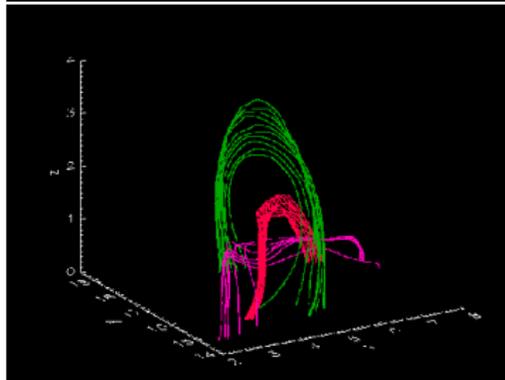
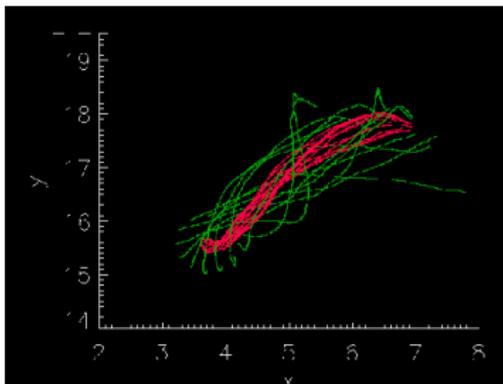
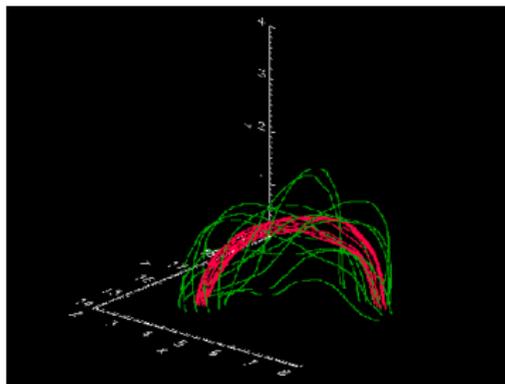
## Post-flare loop in the Bastille event



Flux rope with estimated twist of about  $2\pi$ .

The loop in figure intersects the photosphere on a circle of diameter equal to the magnetogram resolution ( $\Delta = 0.2$ ).

## Post-flare loop in the Bastille event



## Post-flare loop in the Bastille event

The extrapolation provides with the 3D magnetic field:

- Helicity, energy
- Topology, current sheets, stability analysis
- Does the flux rope exist prior the CME, and is it the only one?  
⇒ extrapolations of time series (in progress)
- Energy and helicity budgets
- Full MHD simulation of extrapolated field: desktop CME (Tibor)?

All very interesting,

**BUT**

can we trust the results of the extrapolation code?

# Outline

- 1 The MF method as an extrapolation tool
  - Magneto-frictional relaxation
  - MF used as an extrapolation technique
- 2 Quality of extrapolations
  - Standard test: Low and Lou
  - Flux rope with return current
  - Flux rope without return current
  - Filament and arcade
- 3 Outlook
  - Mgm preprocessing

## Extrapolation of photospheric measurements

- Find the magnetic field in a numerical box for given conditions at the photospheric boundary, assuming a perfectly force-free coronal field.
- Discard non-force-free effects close to the photosphere and errors in, or inconsistencies of, measured photospheric fields.

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## Frictional relaxation for MHD equilibrium solutions

The MF relaxation seeks for a solution of the time-independent MHD with prescribed boundary conditions introducing

- a time-like variable (ie using time-dependent MHD);
- a fictitious friction  $\nu(x, y, z)$  in the momentum balance equation.

Starting from an approximate solution  $\implies$  drive the system toward equilibrium under the combined effect of dissipation and boundary conditions.

## Force-free approximation

For a force-free field, the momentum balance equation,

$$\vec{v} = \frac{1}{\nu} \vec{J} \times \vec{B},$$

gives the velocity field that drives the system toward the force-free equilibrium in terms of the magnetic field.

Hence, only

$$\frac{\partial \vec{B}}{\partial t} = \vec{\nabla} \times (\vec{v} \times \vec{B})$$

need be advanced in time  $\implies$  reduced numerical effort.

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## Initial and boundary conditions

What turns the MF relaxation into an extrapolation technique are

- Initial condition: from the normal component of the vector magnetogram the potential field is computed (Seehafer).
- Boundary conditions: the whole vector magnetogram is copied into the photospheric layer prior to relaxation, and then kept fixed.

The MF relaxation is then started, seeking the equilibrium solution that is force-free, has the magnetogram as (photospheric) boundary, and has small  $\vec{\nabla} \cdot \vec{B}$  errors.

## Boundary conditions

Two ghost layers of bc for  $\vec{B}$  are prescribed by imposing

- Side and top boundaries (both ghost layers):

$$\text{normal component : } \vec{\nabla} \cdot \vec{B} = 0$$

$$\text{transverse components: } \vec{J} \times \vec{B} = 0$$

- Photospheric boundary

- inner layer:

all components: magnetogram

- outer layer:

$$\text{normal component : } \vec{\nabla} \cdot \vec{B} = 0$$

transverse components: : 4th polynomial

The velocity field is windowed toward all but the photospheric boundary.

Flexibility: different bc can be applied to improve relaxation.

## Performance tests

The extrapolation code was tested using four classes of nlfff

- Low and Lou
- Török and Kliem: model of coronal loop with return current
- Titov Demoulin: model of coronal loop with net current
- van Ballegooijen: filament and arcade

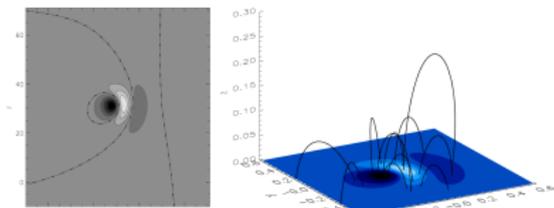
In all cases, the three magnetogram components at  $z = 0$  are the only information used in the extrapolation.

# Outline

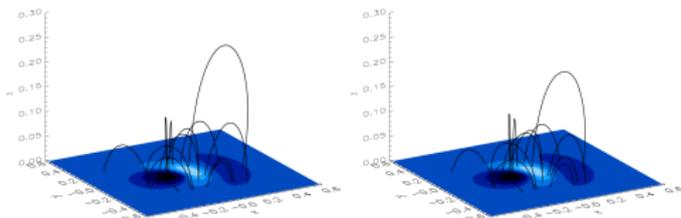
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## The Low and Lou test case

Semi-analytical solution to the nonlinear force-free equations<sup>1</sup>.



Magnetogram ( $B_z$ ) and initial (potential) field.



Input and extrapolated force-free solutions.

$$\sigma_J = \langle \sin \widehat{\vec{J} \cdot \vec{B}} \rangle_J = 0.015, \quad \max |\vec{\nabla} \cdot \vec{B}| = 3 \times 10^{-7}$$

Accurate reconstruction: force and divergence-free

<sup>1</sup>Low & Lou, ApJ 352 343 (1990)

## Comparison with other methods

Several metrics can be defined to compare input  $\vec{b}$  and extrapolated  $\vec{B}$  fields<sup>2</sup>, e.g.

$$E'_M \equiv 1 - E_M; \quad E_M \equiv \frac{1}{N} \sum_i \frac{|\vec{b}_i - \vec{B}_i|}{|\vec{B}_i|},$$

|    | $C_{vec}$ | $C_{CS}$ | $E'_N$ | $E'_M$ | $\epsilon$ | $L_f$ | $L_d$      | $\sigma_J \times 10^2$ |
|----|-----------|----------|--------|--------|------------|-------|------------|------------------------|
| MF | 1.00      | 0.95     | 0.96   | 0.75   | 1.01       | 23.   | $10^{-06}$ | 1.50                   |
| W  | 1.00      | 0.91     | 0.92   | 0.66   | 1.04       | 524.  | 205.       | 4.49                   |

Figures of merit for the extrapolation of Case II, inner domain, with MF and Wiegelmann's (W) codes.

⇒ In the LL case, the MF method gives the most accurate reconstruction among several methods (optimisation, Grad-Rubin, integral).

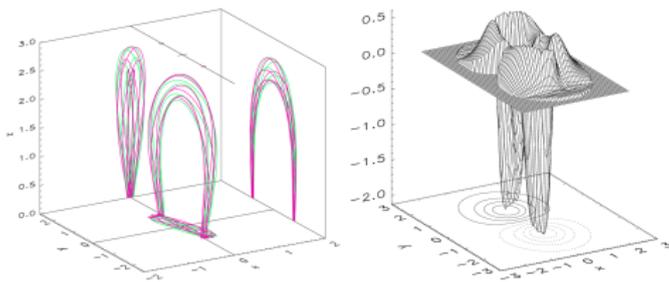
<sup>2</sup>Schrijver *et al*, Sol.Phys. **235** 161 (2006)

# Outline

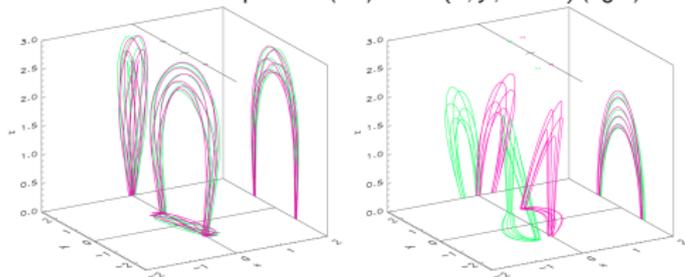
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## Nonlinear dipolar field: TK

Loop with return current and a twist that is about 80% of the instability threshold<sup>3</sup>.



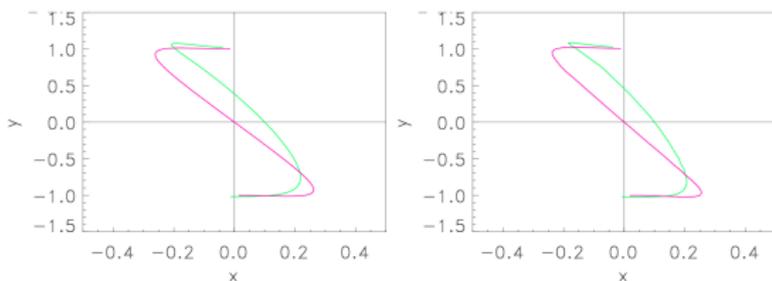
Field lines of the input field (left) and  $\alpha(x, y, z = 0)$  (right)



Extrapolated fields: nonlinear (left) and linear for the best fitting  $\alpha = 0.5$  (right)

# Nonlinear dipolar field: TK

Twist is excellently reproduced



Projections of the **central** field line and a **twisting line** for the input (left) and extrapolated (right) fields

The nonlinear extrapolation is successful in reproducing both the global structure and the details of the twisted loop ( $\sigma_J = 0.01$ ).

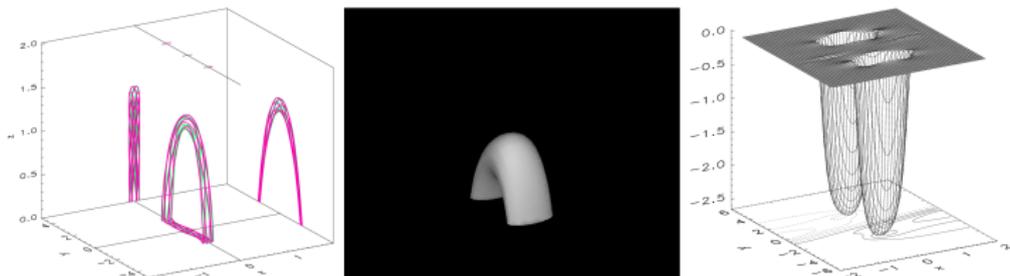
The linear extrapolation fails in both aspects: two flux tubes instead of one, and the best fitting value  $\alpha = 0.5$  is much smaller than the actual maximum value of  $|\alpha(z = 0)|$ .

# Outline

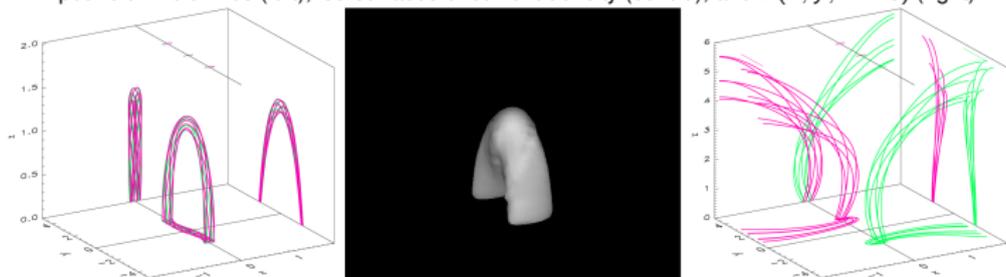
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## Nonlinear quadrupolar field: TD

Loop with no return current and a twist that is about 60% of the kink instability threshold<sup>4</sup>.



Input field: Field lines (left), iso-surfaces of current density (centre), and  $\alpha(x, y, z = 0)$  (right)

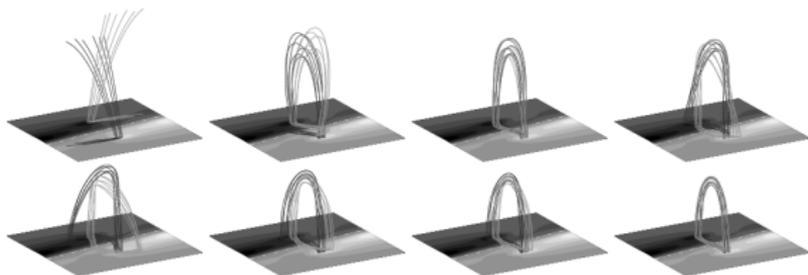


Extrapolated: nonlinear (left), iso-surfaces of current density (centre), and linear for the best fitting  $\alpha = -0.85$  (right)

<sup>4</sup> Kliem and Török, 2004, A&A 413 L23

## Nonlinear quadrupolar field: TD

A very specific balance between the Lorentz self-force of the current-carrying flux rope and the force connected with the external potential field insures a stable equilibrium. 🦋



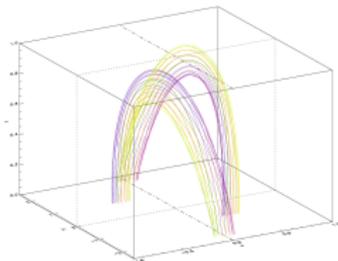
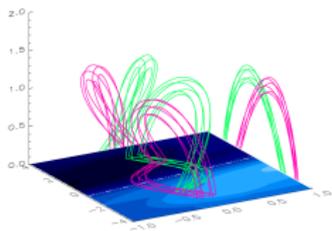
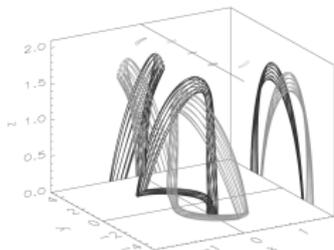
From left to right and top to bottom: evolution at  $(0, 5, 10, 15, 45, 90, 110, 230) \times 10^3$  iterations

The nonlinear extrapolation successfully reproduces the complexity of the original field in all its aspects ( $\sigma_J = 0.02$ ).

In this case, the linear extrapolation fails even more dramatically than in the dipolar case.

# TD puzzle: Imperfect reconstruction

📄: let's raise the upper boundary a bit ... 📄  
Method failure or peculiar equilibrium?



Extrapolation of TD using MF, Optimisation<sup>5</sup>, and full MHD codes

Testing of boundary conditions (Tibor): flexibility is essential.

<sup>5</sup>Wiegelmann *et al*, 2006 A&A, 453, 737-74

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# Aad van Ballegooijen's field

## *Non-linear force-free field (NLFFF) modeling: applications to solar-like data*

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Graham Barnes<sup>b</sup>, Bernhard Kliem<sup>c</sup>, Yang Liu<sup>d</sup>, Jim McTiernan<sup>e</sup>,  
Zoran Mikic<sup>f</sup>, Stéphane Régnier<sup>g</sup>, Slava Titov<sup>h</sup>, Gherardo Valori<sup>c</sup>,  
Aad van Ballegooijen<sup>b</sup>, Brian Welsch<sup>a</sup>, Mike Wheatland<sup>i</sup>,  
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a) Lockheed Martin Advanced Technology Center

b) Northwest Research Associates,  
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c) Astrophysical Institut Potsdam

d) Stanford University

e) University of California, Berkeley

a) Science Applications International Corporation

b) University of St Andrews

c) Harvard-Smithsonian Center for Astrophysics

d) University of Sydney

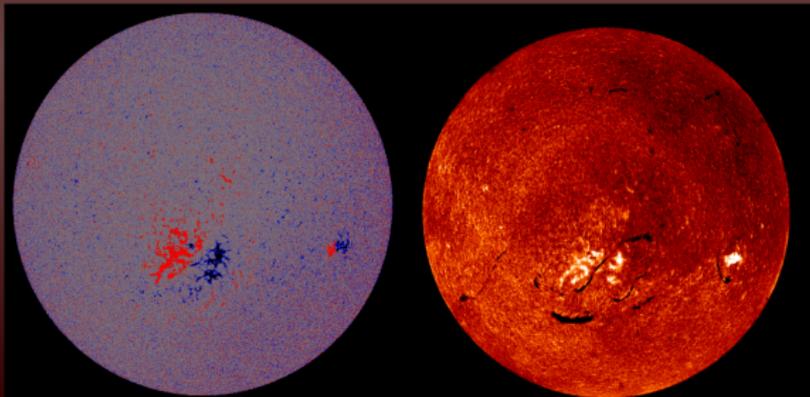
e) Max-Planck-Institut für Solarsystemforschung

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## Aad van Ballegooijen's field

### *"Flux Rope" Case from van Ballegooijen*

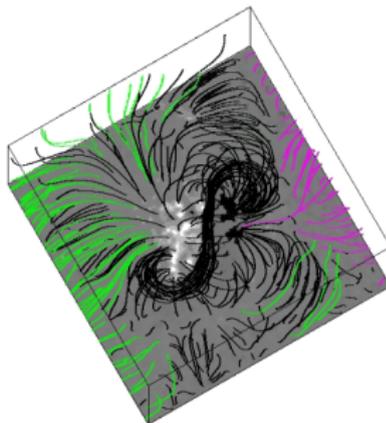
- The MDI magnetogram from 2005 Oct 10 shows a large, reasonably isolated active region in the southern hemisphere.
- The corresponding H $\alpha$  image shows a filament along the polarity inversion line.



# Aad van Ballegooijen's field

## *"Flux Rope" Case from van Ballegooijen*

- Aad took the potential-field extrapolation for this magnetogram, and inserted an S-shaped flux rope in the domain at the location of the filament.
- His magnetofrictional code then relaxed the field to a force-free state.



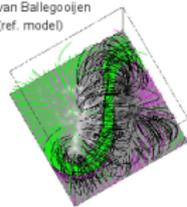
6

# Aad van Ballegooijen's field

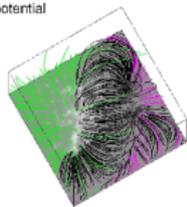
Preliminary results

## *"Flux Rope" Case: Results*

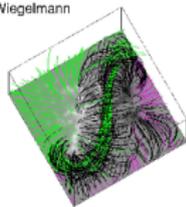
van Ballegooijen  
(ref. model)



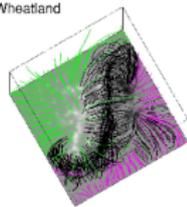
potential



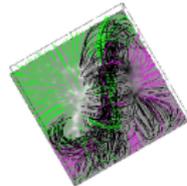
Wiegmann



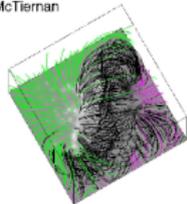
Wheatland



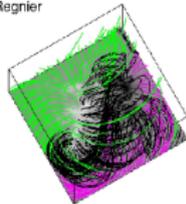
Valori



McTiernan



Regnier



# Aad van Ballegooijen's field

Ours:  $\sigma_J = 0.1$  mgm force-free compatible, but small scales are present

## Quantitative Comparison Metrics

|                    | <u>CS</u> | <u>1-NVE</u> | <u>CWcos</u> | <u>E/Epot</u> |
|--------------------|-----------|--------------|--------------|---------------|
| reference model    | 1.00      | 1.00         | 0.98         | 1.36          |
| Wiegelmann .....   | 0.99      | 0.93         | 0.96         | 1.34          |
| Wheatland .....    | 0.93      | 0.77         | 0.97         | 1.21          |
| Valori .....       | 0.92      | 0.75         | 0.92         | 1.29          |
| McTiernan .....    | 0.89      | 0.71         | 0.84         | 1.05          |
| Régnier* .....     | 0.90      | 0.59         | 0.69         | 1.34          |
| potential solution | 0.90      | 0.65         | (undef.)     | 1.00          |

\*Régnier's calculation was performed at lower resolution.

CS = Cauchy-Schwarz

$$CS \equiv \frac{1}{N} \sum_i \frac{|\vec{v}_i \cdot \vec{w}_i|}{|\vec{v}_i| |\vec{w}_i|}$$

NVE = Normalized  
Vector Error

$$NVE \equiv \frac{\sum_i |\vec{w}_i - \vec{v}_i|}{\sum_i |\vec{v}_i|}$$

$$CW_{\cos} = \frac{\sum_i |\vec{J}_i| \sigma_i}{\sum_i |\vec{J}_i|}$$

$$\sigma_i = \frac{|\vec{J}_i \cdot \vec{B}_i|}{|\vec{J}_i| |\vec{B}_i|}$$

CWcos =  
current-weighted  
cosine between  
**J** and **B**

# Aad van Ballegooijen's field

## *Concluding Remarks*

- All algorithms yield NLFFFs that agree qualitatively well with the reference field, especially in finding the horizontal field structure underneath the arcade.
- Free energy estimates came within 30%, and the best was within 1%.
- Typical runtimes for a  $200^3$  box are of order 10 CPU-hr for the two fastest algorithms.
- The application of these methods to photospheric vector fields look promising.

## Summary

- As long as magnetograms are ff-consistent and scales well resolved, the MF method reproduces the field with high accuracy
- The method and its implementation are very flexible in allowing for
  - non force-free effects
  - tailoring of non-photospheric boundary conditions
  - different discretisation
  - treatment of large dataset using stretched grids
- Reconstruction quality seems to be influenced by the presence of small length scales in the mgm, but not by flux balance, extension of nonlinearities, fl connectivity

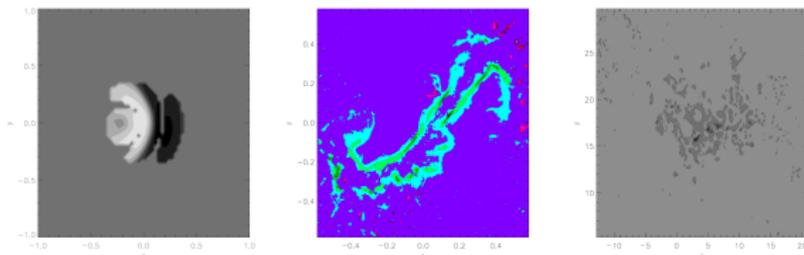
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## Length scales and force-freeness

Small length scales in the mgm  $\implies$  worse relaxation (and reconstruction)

$\alpha$  contour plots for the Low and Lou, van Ballegoijen, and Bastille cases

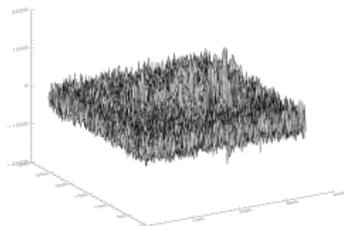


$\sigma_J =$

0.01

0.1

$>0.14$



In a measured magnetogram

- small scales
- error in 180 deg ambiguity removal
- noise
- finite  $\beta$

# Preprocessing

Magnetograms can be preprocessed prior to extrapolation in order to

- reduce noise and smooth small scales  
⇒ facilitate extrapolation
- remove the non force-free part  
⇒ study influence of the (thin) non-force-free layer

Preprocessing as an investigation tool.

## Preprocessing equations

It is based on a minimisation process that changes the field within measurement errors.

From the force-free condition,  $\vec{\nabla} \cdot (\vec{B}\vec{B} - 1/2B^2)$ , applied to the  $mgm^7$

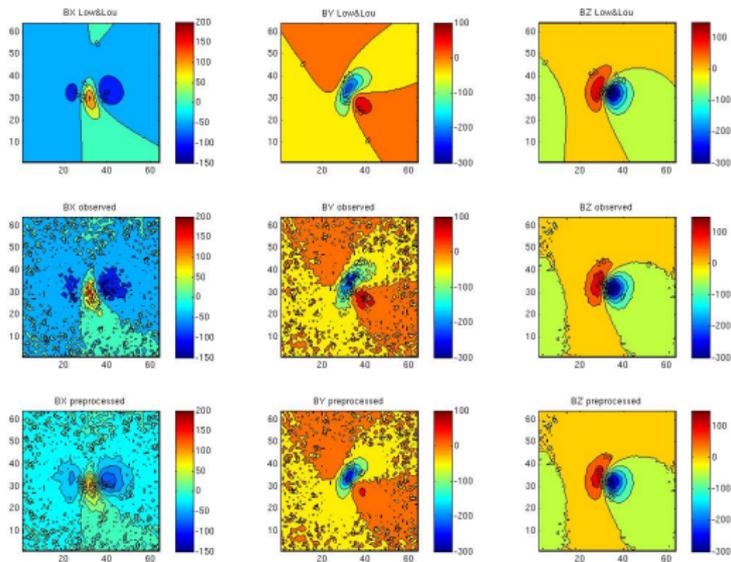
$$L_{force} = \int_{mgm} d\vec{x} \left[ (B_x B_z)^2 + (B_y B_z)^2 + (B_z^2 - B_x^2 - B_y^2) \right] = 0$$

and analogous expression for the torque.

Simulated annealing is used to find the magnetic field that minimises L.

Similarly, a smoothing operator can be devised, but so far results are unsatisfactory.

## Preprocessing example: Low and Lou



The difference in L between "observed" and preprocessed magnetograms is 3 to 5 orders of magnitude.

⇒ Easy to reduce to force-free, less to smooth without flattening excessively the field profiles.

## What we need

- Improve discretisation
- Which techniques to remove noise?
- And small scales?