High Resolution Spectroscopy of Active Galactic Nuclei

Andy Young

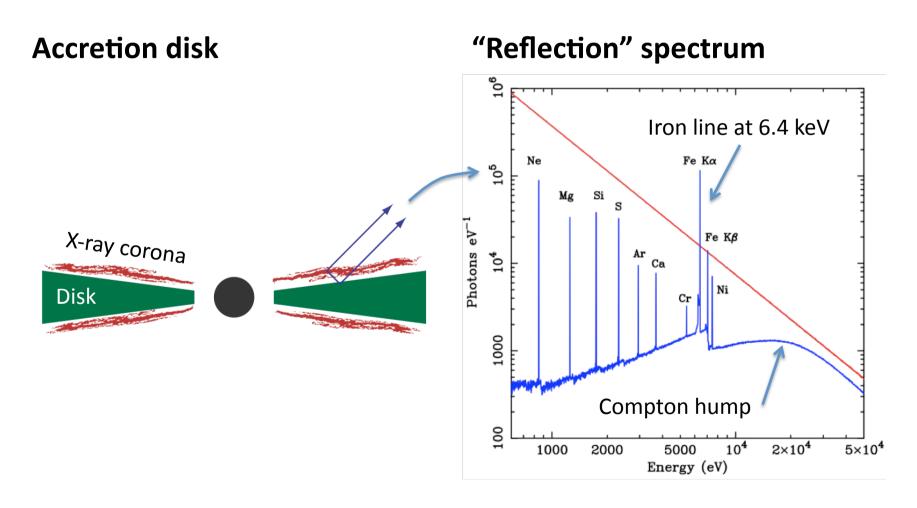


Outline

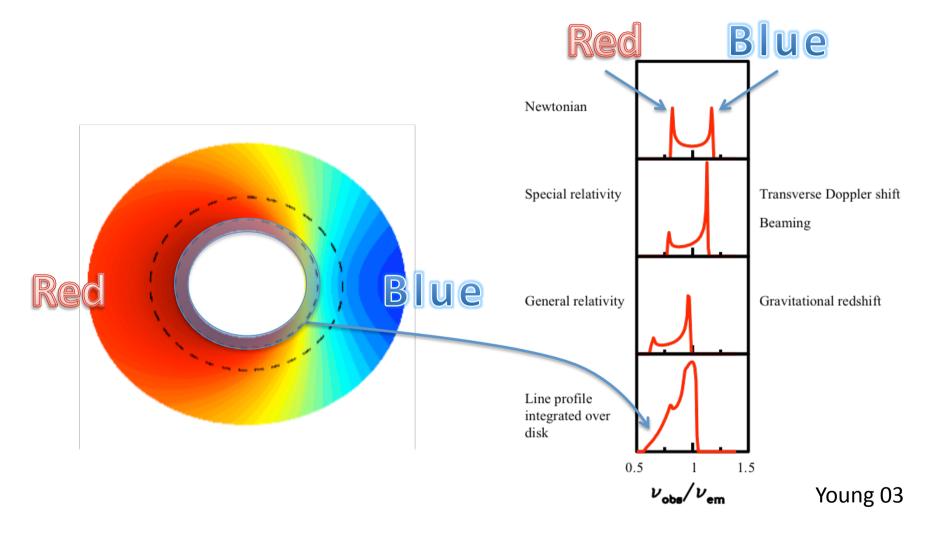
- Warm absorbers and broad iron lines
- Radiatively inefficient accretion flows
- "Towards IXO"

Warm absorbers and broad iron lines

Iron Kα fluorescence line

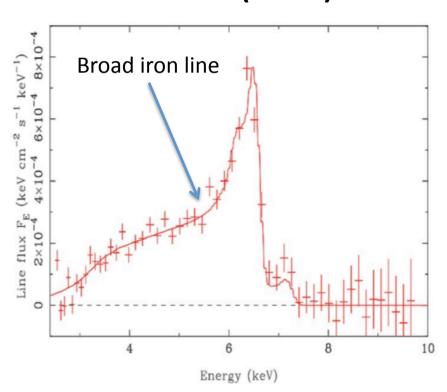


Broad iron line

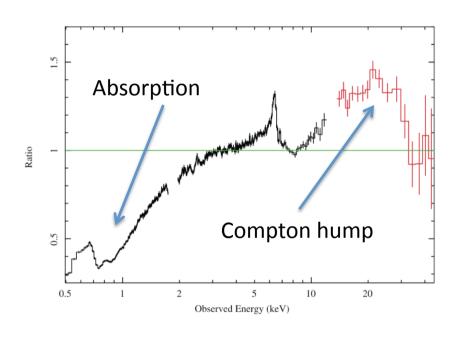


Observations

MCG-6-30-15 (XMM)



MCG-6-30-15 (Suzaku)



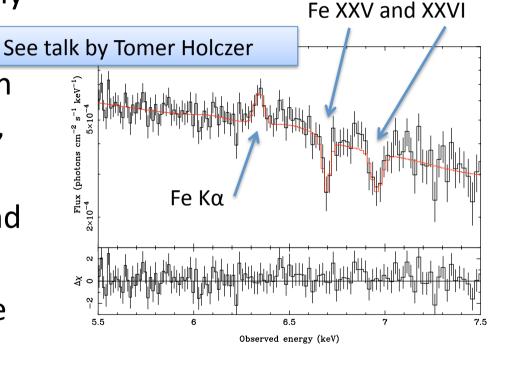
Fabian+02 Miniutti+07

High resolution spectroscopy: MCG—6-30-15 (Chandra HETGS)

Ionized absorber is clearly present and important

 Broad range of ionization states, column densities, outflow velocities

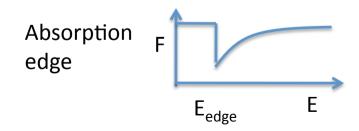
- Absorption by Fe XXV and Fe XXVI
- What does this do to the hard X-ray continuum?

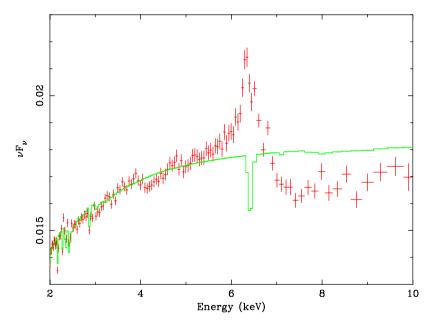


Young+05

Ionized absorption

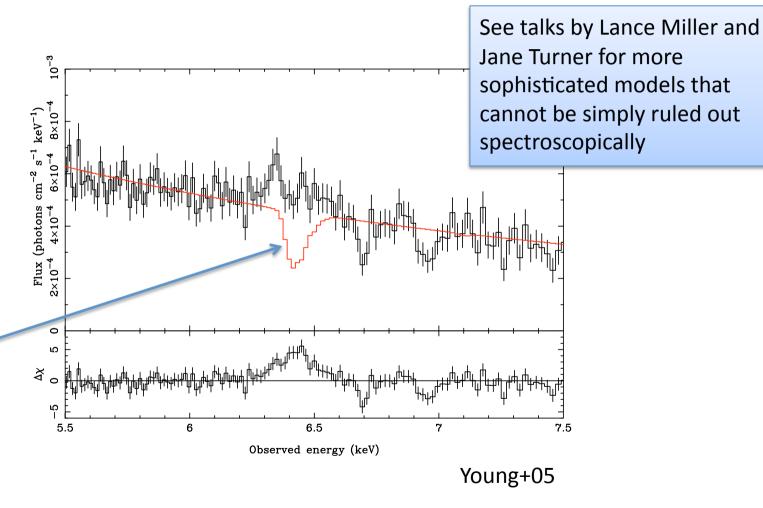
- Could ionized absorption extend up to the Fe K band?
- Does this affect our interpretation of the broad iron line?





Cartoon of simple ionized absorption model

Can rule out the *simplest* ionized absorption models



These

absorption

predicted by

lines are

all simple

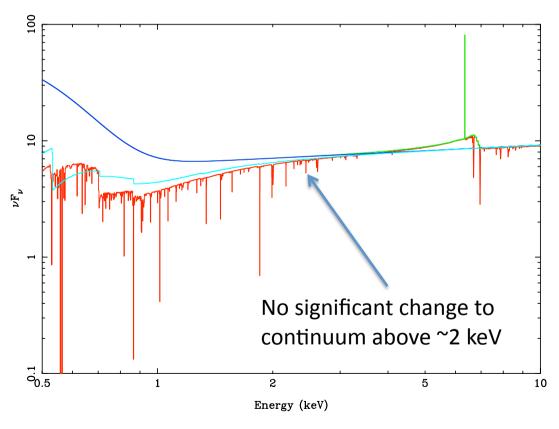
absorption

ionized

models

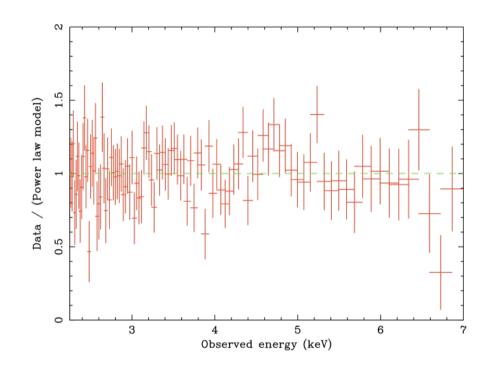
Broad iron line still robust

The best-fitting warm absorber model still *requires* a broad iron line



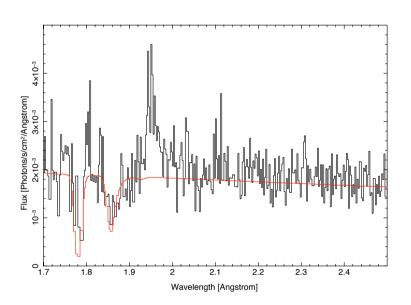
Variability

- Difference spectrum of high—low states well described by a powerlaw
- Also suggests broad
 Fe K feature is not the
 result of absorption



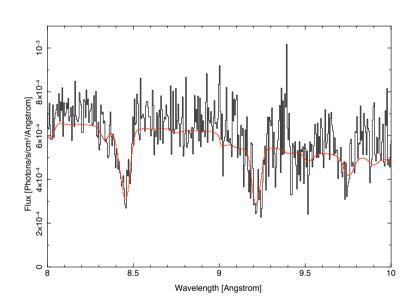
Other AGN, e.g., NGC 3516...

Fe K bandpass



8-10 A

See talk by Jane Turner and poster by Missagh Mehdipour



Need detailed models of ionized absorbers in different spectral sates and to consider the hard X-ray spectrum and broad iron line. Self-consistent physical model of warm absorber and broad iron line. Work in progress...

Momtahan+ (in prep)

Towards IXO...

- High resolution spectroscopy up to Fe K band
 - Measure properties of absorber line strengths, widths, velocities
- Broad bandpass
 - Simultaneously measure continuum across broad X-ray bandpass
- Large area
 - Time- and flux-resolved spectroscopy
 - Routinely allow detailed studies of fainter AGN beyond the usual suspects
- Also need self-consistent physical models of these systems